6.869 Advances in Computer Vision Bill Freeman and Antonio Torralba

Lecture 12 Calibration and Stereo

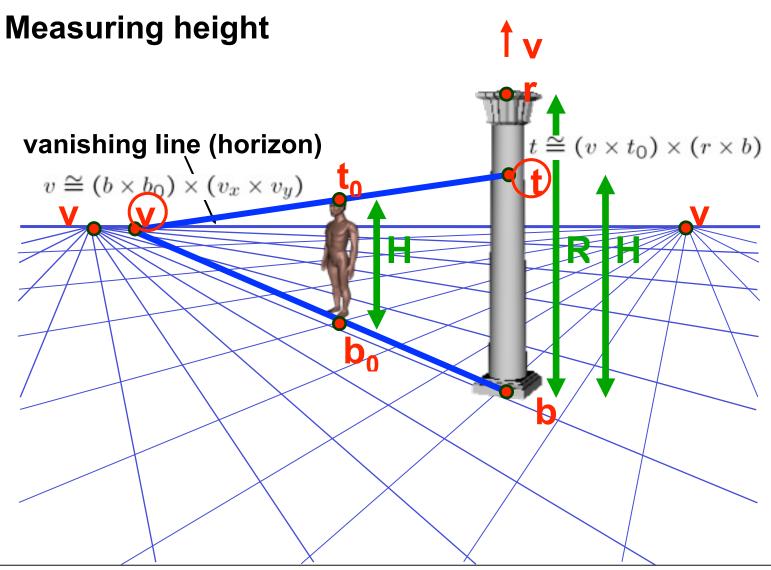
Szeliski book: section 2.1, section 7.2, chapter 11.



Camera calibration

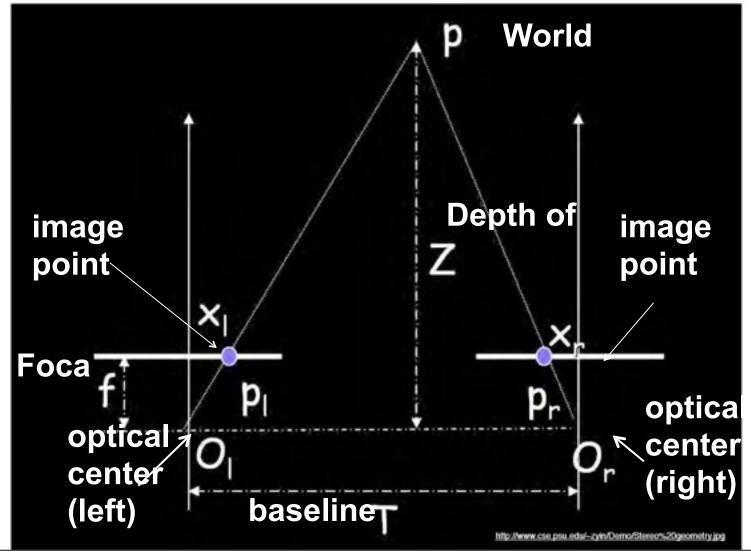
- Use the camera to tell you things about the world:
 - Relationship between coordinates in the world and coordinates in the image: *geometric camera calibration, see* Szeliski, chapter 6 (see ch. 11 on stereo)

One reason to calibrate a camera



Monday, March 14, 2011

Another reason to calibrate a camera



Monday, March 14, 2011

Three camera projections

3-d point 2-d image position (1) Perspective: $(x, y, z) \rightarrow \left(\frac{fx}{z}, \frac{fy}{z}\right)$ $(x, y, z) \rightarrow \left(\frac{fx}{z_0}, \frac{fy}{z_0}\right)$ (2) Weak perspective: (3) Orthographic: $(x, y, z) \rightarrow (x, y)$

Is the perspective projection a linear transformation?

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no-division by z is nonlinear

Homogeneous coordinates

Is the perspective projection a linear transformation?

no-division by z is nonlinear

Homogeneous coordinates

Is the perspective projection a linear transformation?

no—division by z is nonlinear Trick: add one more coordinate:

 $(x, y) \Rightarrow \begin{bmatrix} x \\ y \\ 1 \end{bmatrix} \qquad (x, y, z) \Rightarrow \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix}$ homogeneous image coordinates homogeneous scene coordinates

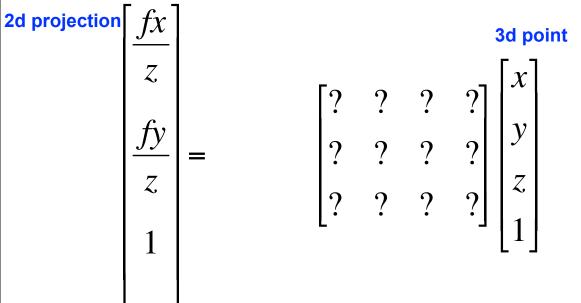
Converting from homogeneous coordinates

$$\begin{bmatrix} x \\ y \\ w \end{bmatrix} \Rightarrow (x/w, y/w) \qquad \begin{bmatrix} x \\ y \\ z \\ w \end{bmatrix} \Rightarrow (x/w, y/w, z/w)$$

2d projection $\begin{pmatrix} fx \\ \overline{z}, \overline{fy} \\ z \end{pmatrix} \stackrel{3d point}{\leftarrow} (x, y, z)$

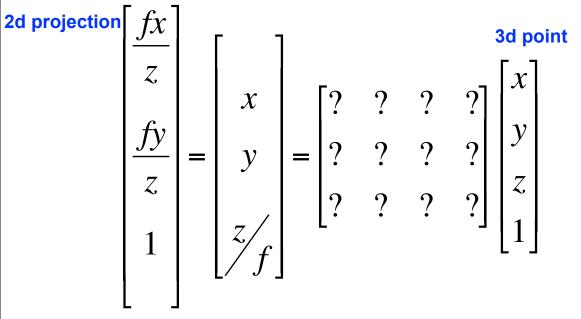
2d projection $\left(\frac{fx}{z}, \frac{fy}{z}\right) \stackrel{\text{3d point}}{\leftarrow} (x, y, z)$

Homogeneous coords



2d projection $\left(\frac{fx}{z}, \frac{fy}{z}\right) \stackrel{\text{3d point}}{\leftarrow} (x, y, z)$

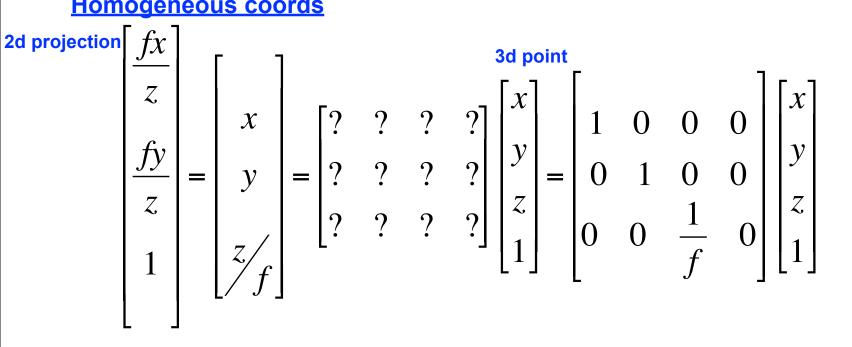
Homogeneous coords



Perspective Projection **Euclidean coords**

2d projection, $\left(\frac{fx}{7}, \frac{fy}{7}\right) \xleftarrow{\text{3d point}}{} (x, y, z)$

Homogeneous coords



2d projection $\left(\frac{fx}{z}, \frac{fy}{z}\right) \stackrel{\text{3d point}}{\leftarrow} (x, y, z)$

End of the second seco

Perspective projection is a matrix multiply using homogeneous coordinates. The matrix is called the projection matrix.

How does scaling the projection matrix change the transformation?

$$\begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 1/f & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix} = \begin{bmatrix} x \\ y \\ z/f \end{bmatrix} \implies \left(f \frac{x}{z}, f \frac{y}{z} \right)$$

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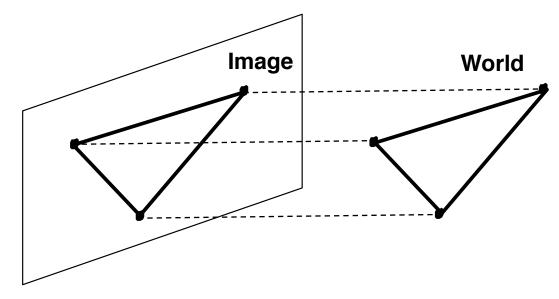
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Special case of perspective projection

• Distance from the COP to the PP is infinite

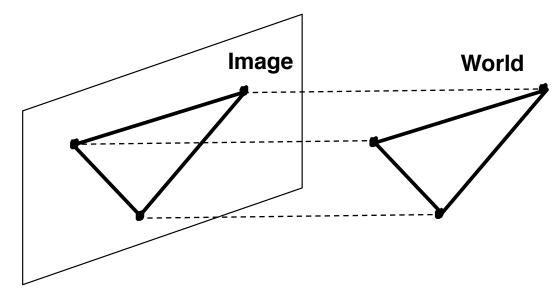


- Also called "parallel projection"
- What's the projection matrix?



Special case of perspective projection

• Distance from the COP to the PP is infinite

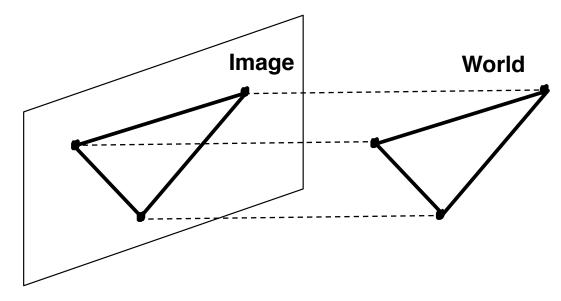


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$$\begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix} = \begin{bmatrix} x \\ y \\ 1 \end{bmatrix} \Rightarrow (x, y)$$

Special case of perspective projection

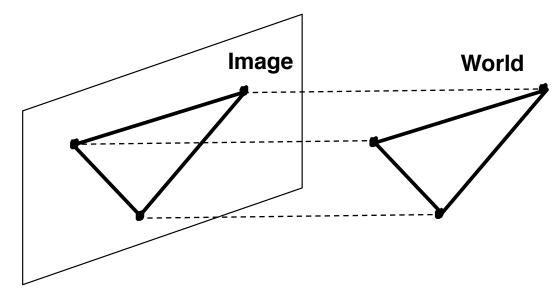
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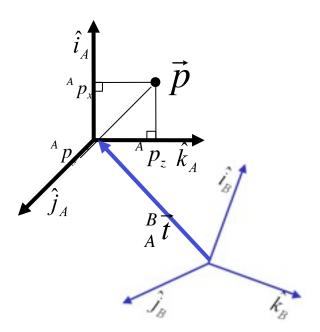
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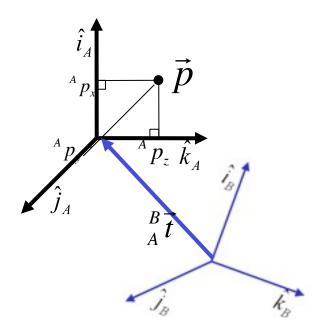


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$$\begin{bmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & 0 & 1 \end{bmatrix} \begin{bmatrix} x \\ y \\ z \\ 1 \end{bmatrix} = \begin{bmatrix} x \\ y \\ 1 \end{bmatrix} \Rightarrow (x, y)$$

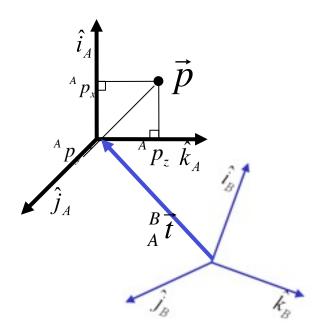






"as described in the coordinates of frame B"

 ${}^{B}\vec{p} = {}^{B}_{A}R {}^{A}\vec{p} + {}^{B}_{A}\vec{t}$

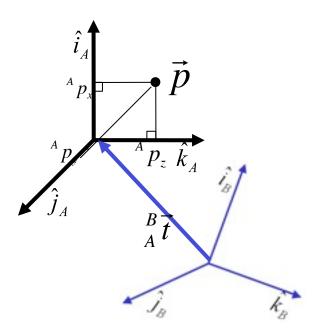


"as described in the coordinates of frame B"

Let's write

 ${}^{B}\vec{p} = {}^{B}_{A}R {}^{A}\vec{p} + {}^{B}_{A}\vec{t}$

as a single matrix equation:

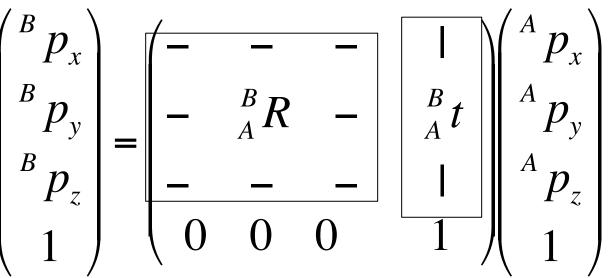


"as described in the coordinates of frame B"

Let's write

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as a single matrix equation:



Α $^{A}p_{z}\hat{k}_{A}$ Translation and rotation, written in each set of coordinates

Non-homogeneous coordinates

$${}^{B}\vec{p} = {}^{B}_{A}R {}^{A}\vec{p} + {}^{B}_{A}\vec{t}$$

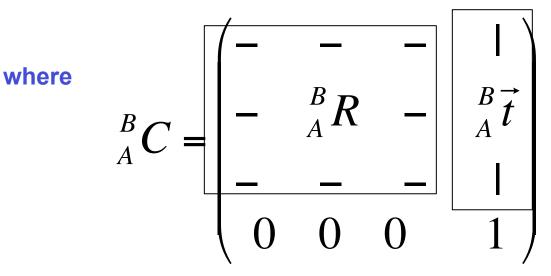
Translation and rotation, written in each set of coordinates

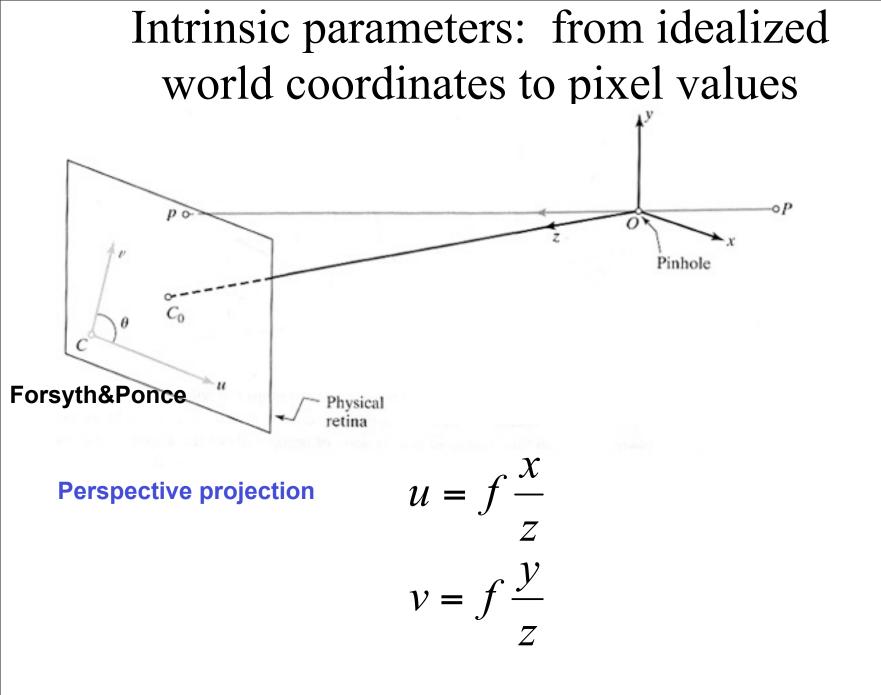
Non-homogeneous coordinates

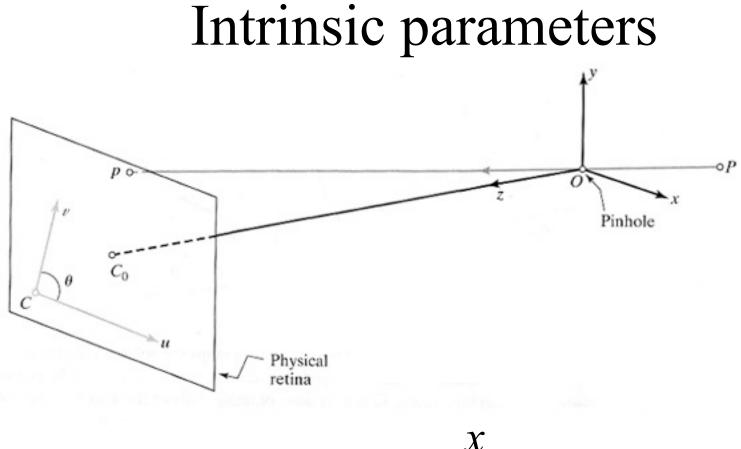
$${}^{B}\vec{p} = {}^{B}_{A}R {}^{A}\vec{p} + {}^{B}_{A}\vec{t}$$

Homogeneous coordinates

$${}^{B}\vec{p}={}^{B}_{A}C{}^{A}\vec{p}$$

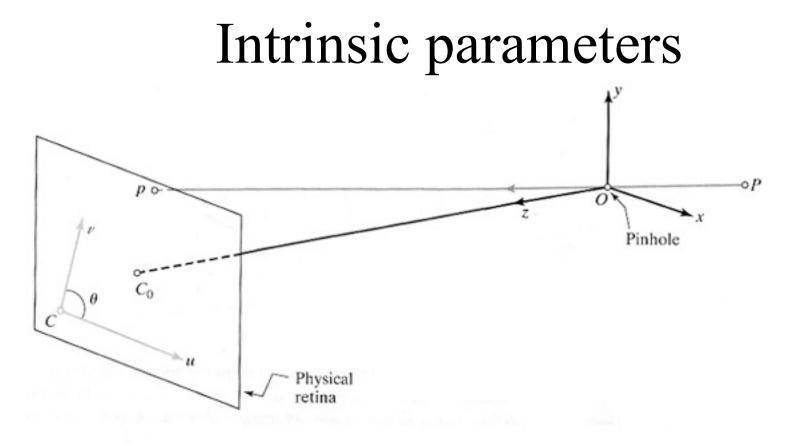






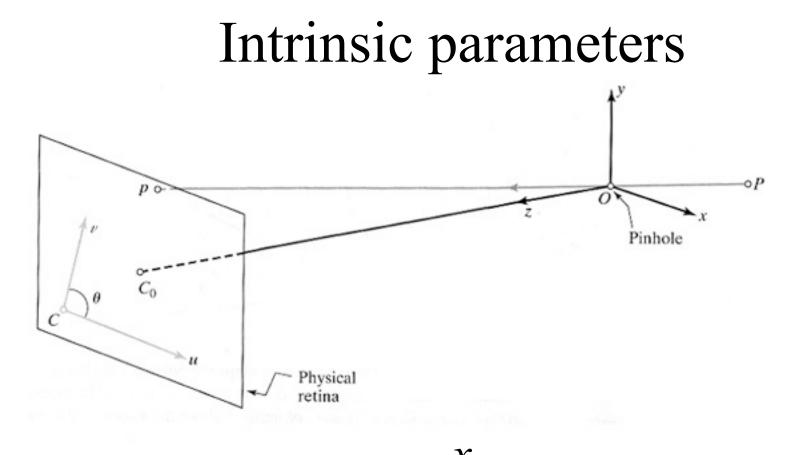
But "pixels" are in some arbitrary spatial units

$$u = \alpha \frac{x}{z}$$
$$v = \alpha \frac{y}{z}$$



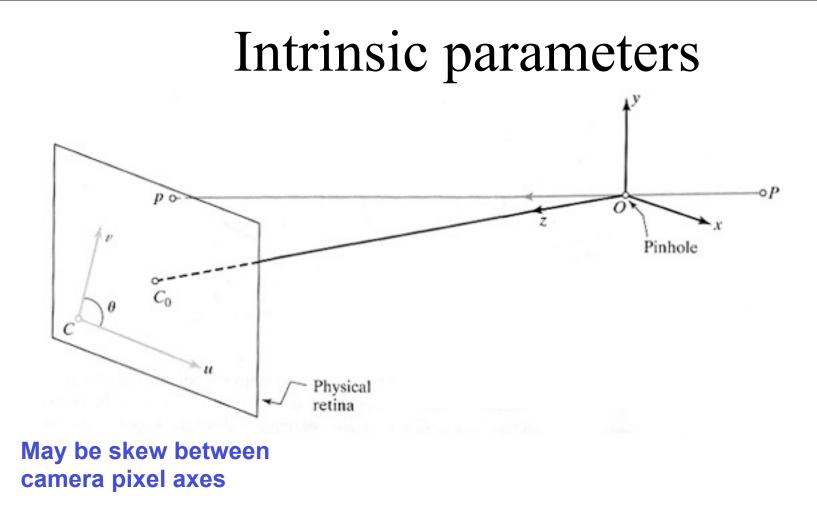
Maybe pixels are not square

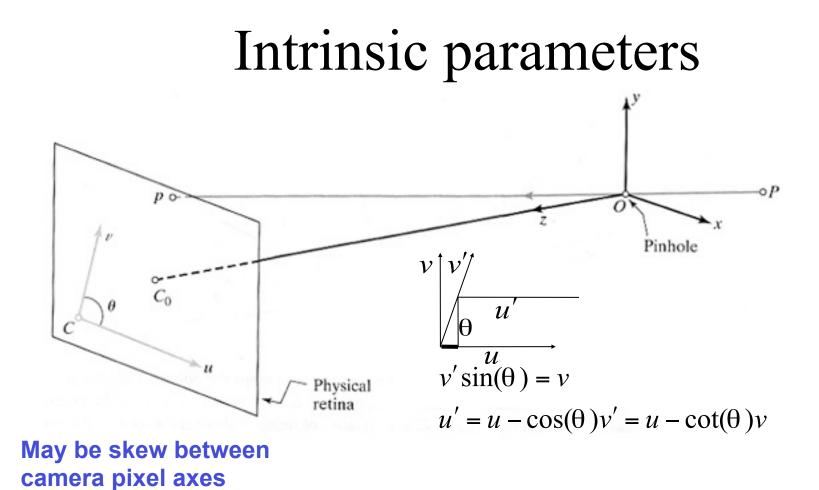
$$u = \alpha \frac{x}{z}$$
$$v = \beta \frac{y}{z}$$



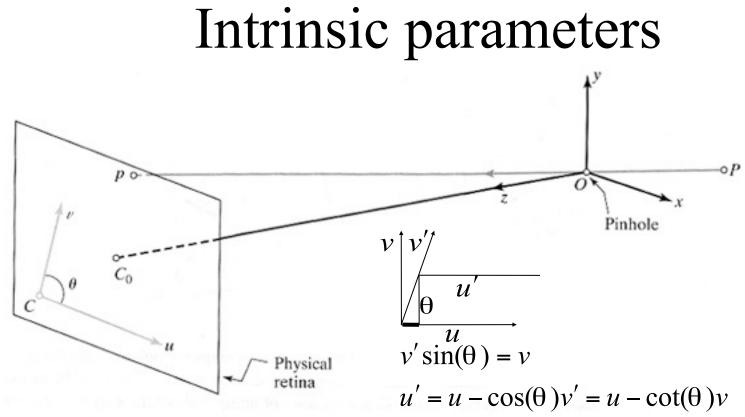
We don't know the origin of our camera pixel coordinates

$$u = \alpha \frac{x}{z} + u_0$$
$$v = \beta \frac{y}{z} + v_0$$



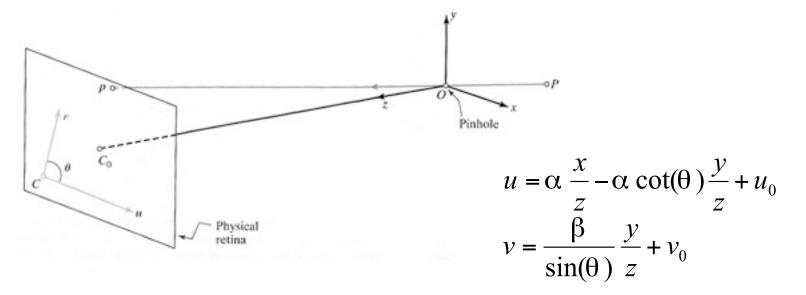


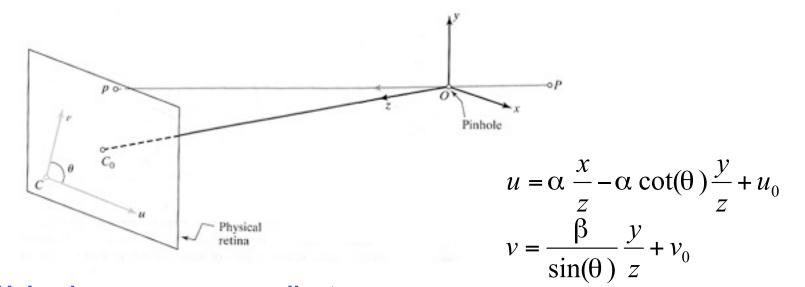
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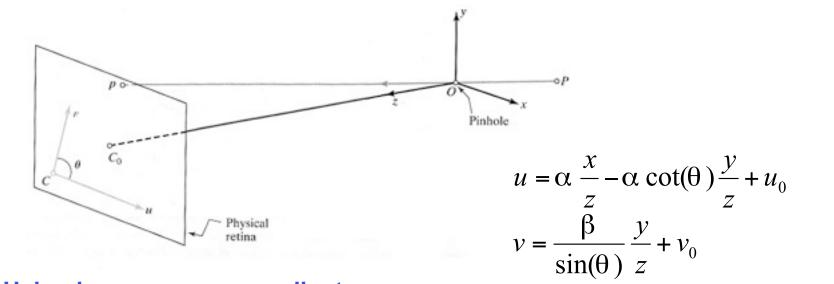
May be skew between camera pixel axes

$$u = \alpha \frac{x}{z} - \alpha \cot(\theta) \frac{y}{z} + u_0$$
$$v = \frac{\beta}{\sin(\theta)} \frac{y}{z} + v_0$$





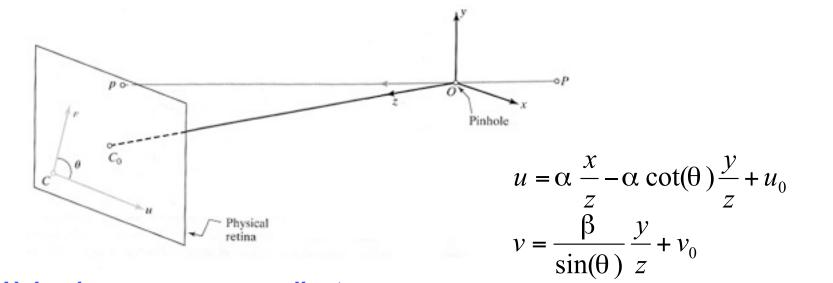
Using homogenous coordinates, we can write this as:



Using homogenous coordinates, we can write this as:

$$\begin{pmatrix} u \\ v \\ 1 \end{pmatrix} = \begin{pmatrix} \alpha & -\alpha \cot(\theta) & u_0 & 0 \\ 0 & \frac{\beta}{\sin(\theta)} & v_0 & 0 \\ 0 & 0 & 1 & 0 \end{pmatrix} \begin{pmatrix} x \\ y \\ z \\ 1 \end{pmatrix}$$

/ \

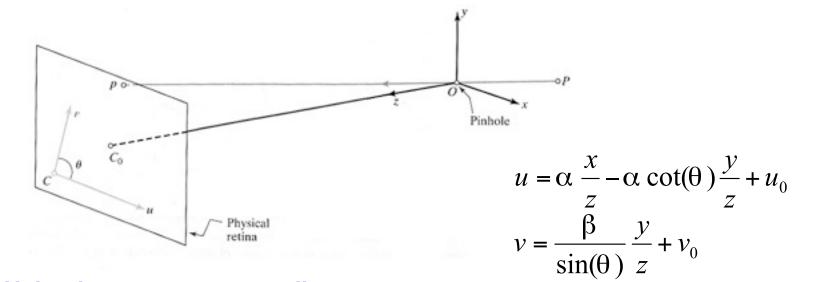


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/ \

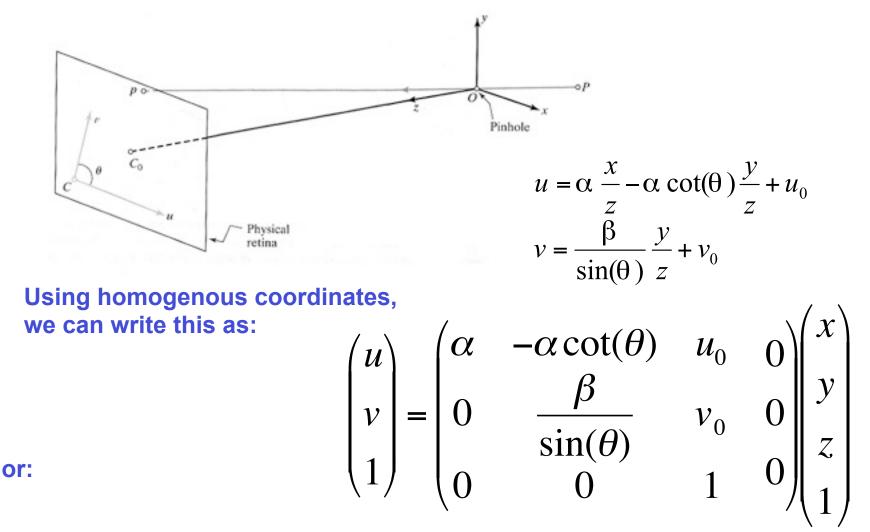
or:



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or:



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In camera-based coords

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In pixels

Extrinsic parameters: translation and rotation of camera frame

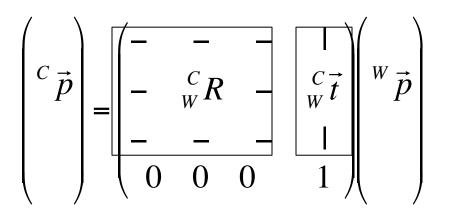
Extrinsic parameters: translation and rotation of camera frame

 ${}^{C}\vec{p} = {}^{C}_{W}R {}^{W}\vec{p} + {}^{C}_{W}\vec{t}$

Non-homogeneous coordinates

Extrinsic parameters: translation and rotation of camera frame

$${}^{C}\vec{p} = {}^{C}_{W}R {}^{W}\vec{p} + {}^{C}_{W}\vec{t}$$



Non-homogeneous coordinates

Homogeneous coordinates

Intrinsic

$$\vec{p} = \mathbf{K} \ \ \vec{p}$$

$${}^{C}\vec{p} = \left[\begin{array}{c} - & - & - \\ - & W R & - \\ - & W R & - \\ - & - & - \\ 0 & 0 & 0 \end{array} \right] \ \mathbf{Extrinsic}$$

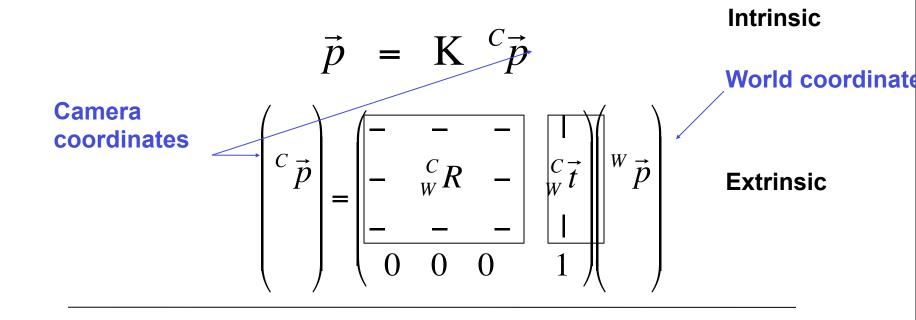
$$\mathbf{Extrinsic}$$

Forsyth&Ponce

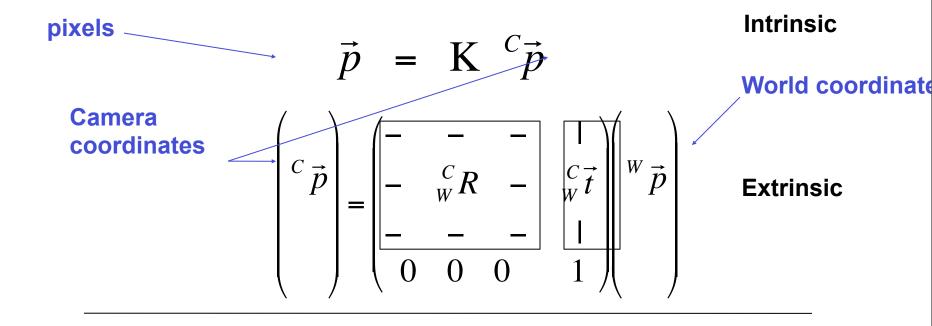
$$\vec{p} = \mathbf{K} \quad \vec{p}$$

$$\vec{p} = \mathbf{K} \quad \vec{p}$$
World coordinate
$$\vec{p} = \begin{pmatrix} - & - & - & - \\ - & W & R & - & - \\ - & W & R & - & - \\ - & 0 & 0 & 0 & 1 \end{pmatrix} \quad \mathbf{W} \quad \vec{p}$$
Extrinsic

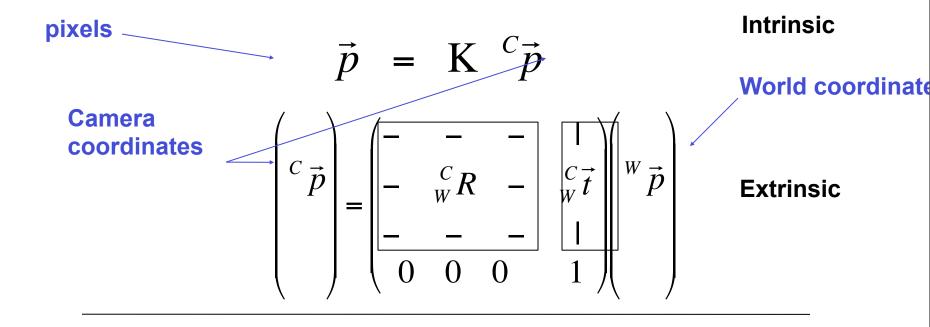
Forsyth&Ponce



Forsyth&Ponce



Forsyth&Ponce



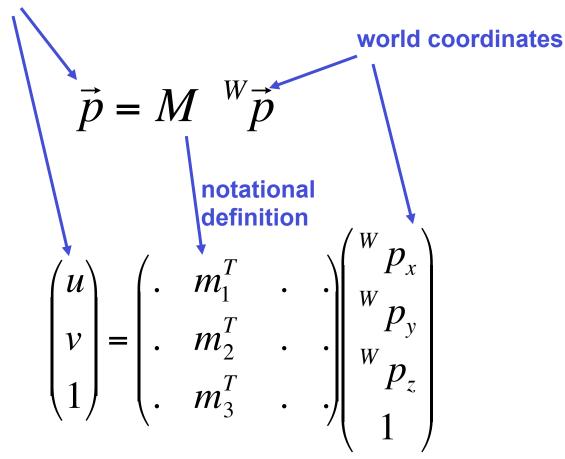
$$\vec{p} = K \begin{pmatrix} C & C & C \\ W & W & W \\ 0 & 0 & 0 \end{pmatrix} \overset{W}{p} \vec{p}$$

$$\vec{p} = M \overset{W}{p}$$

Forsyth&Ponce

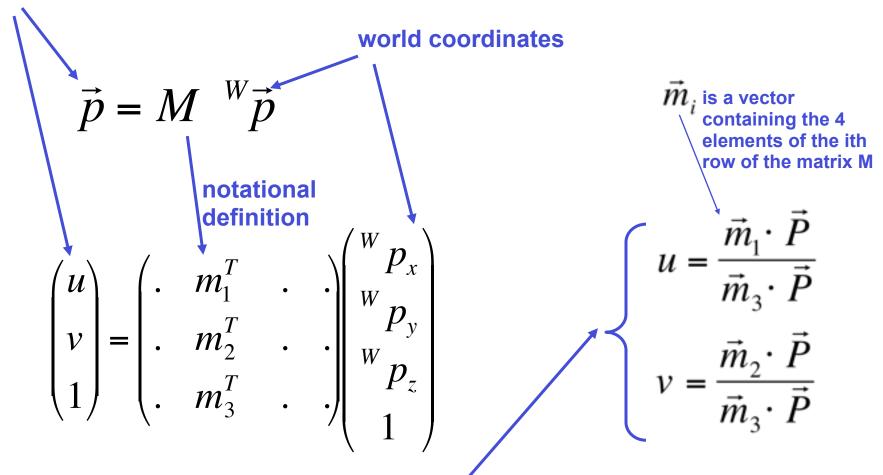
Other ways to write the same equation

pixel coordinates



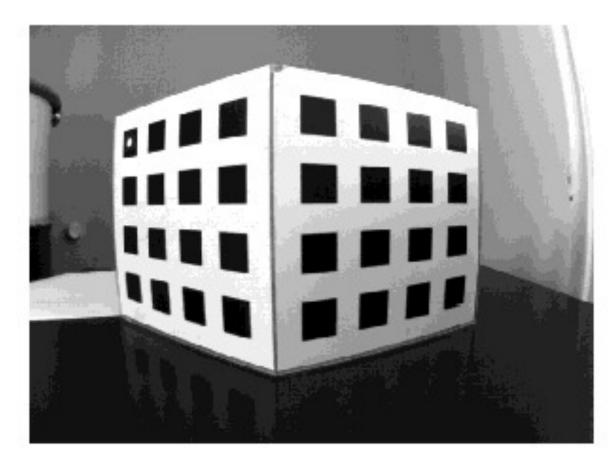
Other ways to write the same equation

pixel coordinates



Conversion of the 2-d position back from homogeneous coordinates leads to:

Calibration target



The Opti-CAL Calibration Target Image Find the position, u_i and v_i , in pixels, of each calibration object feature point.

http://www.kinetic.bc.ca/CompVision/opti-CAL.html

From before, we had these equations relating image positions, u,v, to points at 3-d positions P (in homogeneous coordinates):

$$u = \frac{\vec{m}_1 \cdot \vec{P}}{\vec{m}_3 \cdot \vec{P}}$$
$$v = \frac{\vec{m}_2 \cdot \vec{P}}{\vec{m}_3 \cdot \vec{P}}$$

From before, we had these equations relating image positions, u,v, to points at 3-d positions P (in homogeneous coordinates):

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$$v = \frac{\vec{m}_2 \cdot \vec{P}}{\vec{m}_3 \cdot \vec{P}}$$

So for each feature point, i, we have:

$$(\vec{m}_1 - u_i \vec{m}_3) \cdot \vec{P}_i = 0$$

$$(\vec{m}_2 - v_i \vec{m}_3) \cdot \vec{P}_i = 0$$

Stack all these measurements of i=1...n points $(\vec{m}_1 - u_i \vec{m}_3) \cdot \vec{P}_i = 0$ $(\vec{m}_2 - v_i \vec{m}_3) \cdot \vec{P}_i = 0$

into a big matrix (cluttering vector arrows omitted from P and m):

$$\begin{pmatrix} P_1^T & 0^T & -u_1 P_1^T \\ 0^T & P_1^T & -v_1 P_1^T \\ \cdots & \cdots & \cdots \\ P_n^T & 0^T & -u_n P_n^T \\ 0^T & P_n^T & -v_n P_n^T \end{pmatrix} \begin{pmatrix} m_1 \\ m_2 \\ m_3 \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ \vdots \\ 0 \\ 0 \end{pmatrix}$$

In vector form:
$$\begin{pmatrix} P_{1}^{T} & 0^{T} & -u_{1}P_{1}^{T} \\ 0^{T} & P_{1}^{T} & -v_{1}P_{1}^{T} \\ \cdots & \cdots & \cdots \\ P_{n}^{T} & 0^{T} & -u_{n}P_{n}^{T} \\ 0^{T} & P_{n}^{T} & -v_{n}P_{n}^{T} \end{pmatrix} \begin{pmatrix} m_{1} \\ m_{2} \\ m_{3} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ \vdots \\ 0 \\ 0 \end{pmatrix}$$

$$\begin{pmatrix} P_{1x} & P_{1y} & P_{1z} & 1 & 0 & 0 & 0 & -u_{1}P_{1x} & -u_{1}P_{1y} & -u_{1}P_{1z} & -u_{1} \\ 0 & 0 & 0 & 0 & P_{1x} & P_{1y} & P_{1z} & 1 & -v_{1}P_{1x} & -v_{1}P_{1z} & -v_{1} \\ \dots & \dots & \dots & \dots \\ P_{nx} & P_{ny} & P_{nz} & 1 & 0 & 0 & 0 & -u_{n}P_{nx} & -u_{n}P_{ny} & -u_{n}P_{nz} & -u_{n} \\ 0 & 0 & 0 & 0 & P_{nx} & P_{ny} & P_{nz} & 1 & -v_{n}P_{nx} & -v_{n}P_{ny} & -v_{n}P_{nz} & -v_{n} \\ \end{pmatrix} \begin{pmatrix} m_{11} \\ m_{22} \\ m_{23} \\ m_{34} \\ m_{31} \\ m_{32} \\ m_{33} \\ m_{34} \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ \vdots \\ 0 \\ 0 \end{pmatrix}$$

We want to solve for the unit vector m (the stacked one) that minimizes $|Qm|^2$

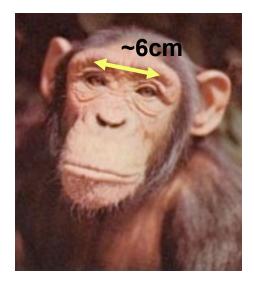
The minimum eigenvector of the matrix $Q^{T}Q$ gives us that (see Forsyth&Ponce, 3.1), because it is the unit vector x that minimizes $x^{T} Q^{T}Q x$.

Once you have the M matrix, can recover the intrinsic and extrinsic parameters.

$$\mathcal{M} = \begin{pmatrix} \alpha \mathbf{r}_1^T - \alpha \cot \theta \mathbf{r}_2^T + u_0 \mathbf{r}_3^T & \alpha t_x - \alpha \cot \theta t_y + u_0 t_z \\ \frac{\beta}{\sin \theta} \mathbf{r}_2^T + v_0 \mathbf{r}_3^T & \frac{\beta}{\sin \theta} t_y + v_0 t_z \\ \mathbf{r}_3^T & t_z \end{pmatrix}$$

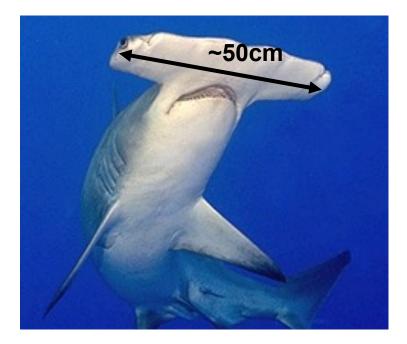
Stereo vision

Stereo vision



Stereo vision

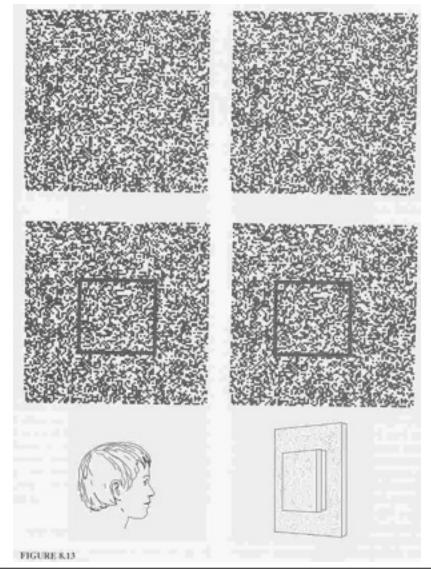




Depth without objects Random dot stereograms (Bela Julesz)

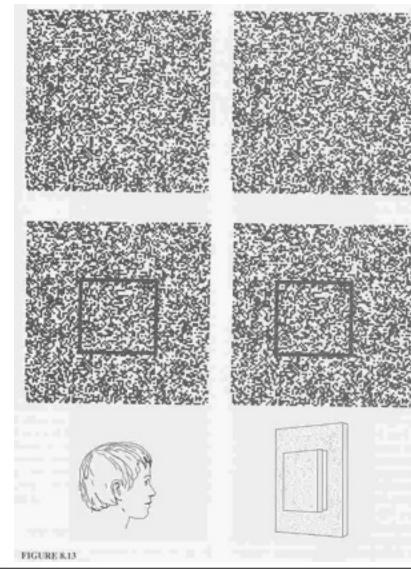


Depth without objects Random dot stereograms (Bela Julesz)





Depth without objects Random dot stereograms (Bela Julesz)



1	0	1	0	1	0	0	1	0	1
1	0	0	1	0	1	0	1	0	0
0	0	1	1	0	1	1	0	1	0
¢	1	0	Y	A	A	8	0	0	1
1	1	1	×	0	4	0	A	0	1
0	0	1	×	A	A	0	A	1	0
1	1	1	Y	0	8	A	8	0	1
1	0	0	1	1	0	1	1	0	1
1	1	0	0	1	1	0	1	1	1
0	1	0	0	0	1	1	1	1	0

1	0	1	0	1	0	ô	1	0	1
1	0	0	1	0	1	0	1	0	0
¢	0	1	1	0	1	1	0	1	0
0	1	0	A	A	8	8	×	٥	1
1	1	1	8	A	8	4	Y	0	1
0	0	1	A		0	A	Y	1	0
1	1	1	0	8	A	8	×	0	1
1	0	Ó	1	1	0	1	1	0	1
1	1	٥	0	1	1	0	1	1	1
0	1	0	0	0	1	1	1	1	0

Julesz, 1971



Depth for familiar objects

http://www.youtube.com/watchv=G_Qwp2GdB1M

(Gregory 1970; Hill and Bruce 1993, 1994; Papathomas and DeCarlo 1999)

Depth for familiar objects



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(Gregory 1970; Hill and Bruce 1993, 1994; Papathomas and DeCarlo 1999)

Stereo photography and stereo viewers

Take two pictures of the same subject from two slightly different viewpoints and display so that each eye sees only one of the images.



Invented by Sir Charles Wheatstone, 1838

Slide credit: Kristen Grauman

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Take two pictures of the same subject from two slightly different viewpoints and display so that each eye sees only one of the images.



Invented by Sir Charles Wheatstone, 1838



Image courtesy of fisher-price.com

Slide credit: Kristen Grauman

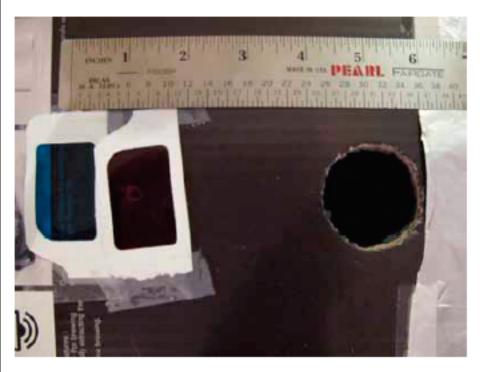


Public Library, Stereoscopic Looking Room, Chicago, by Phillips, 1923



de credit: Kristen Grauman

Anaglyph pinhole camera







Autostereograms

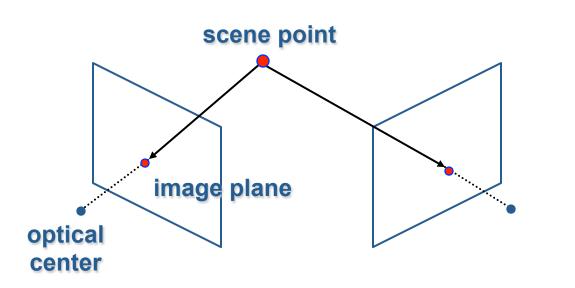


Exploit disparity as depth cue using single image.

(Single image random dot stereogram, Single image stereogram)

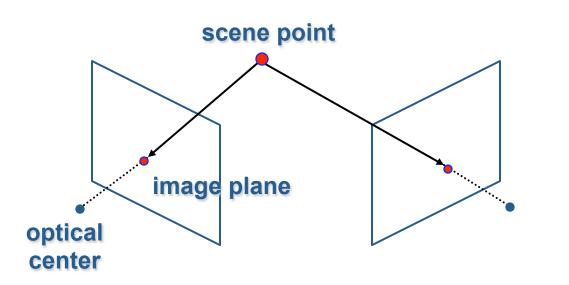
Images from magiceye.com

Slide credit: Kristen Grauman



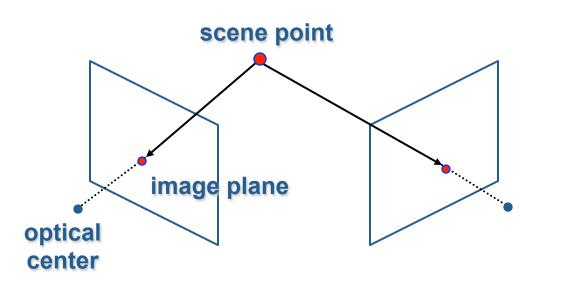


• Stereo: shape from disparities between two views



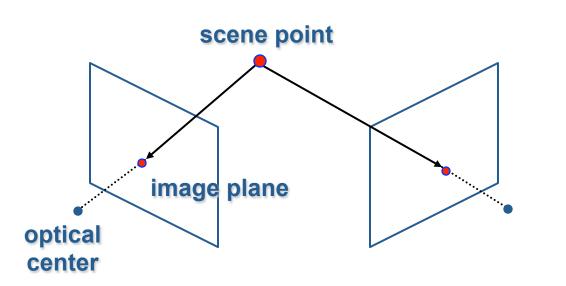


- Stereo: shape from disparities between two views
- We'll need to consider:



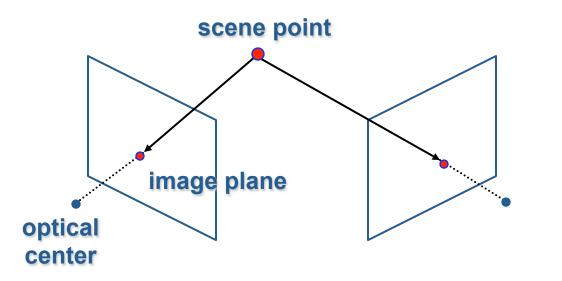


- Stereo: shape from disparities between two views
- We'll need to consider:
- Info on camera pose ("calibration")





- Stereo: shape from disparities between two views
- We'll need to consider:
- Info on camera pose ("calibration")
- Image point correspondences

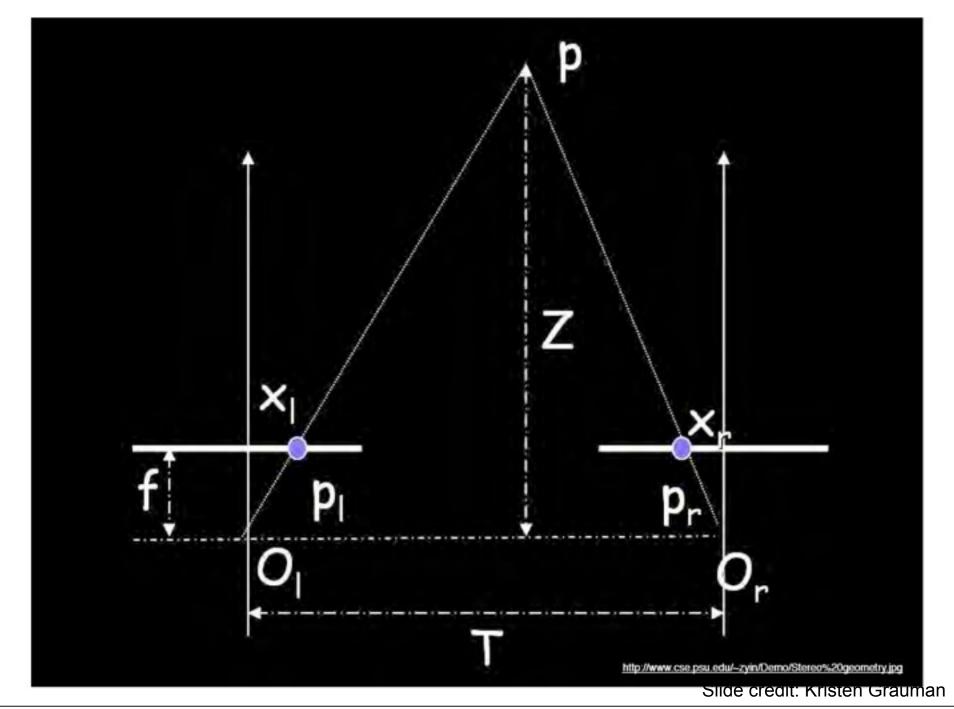




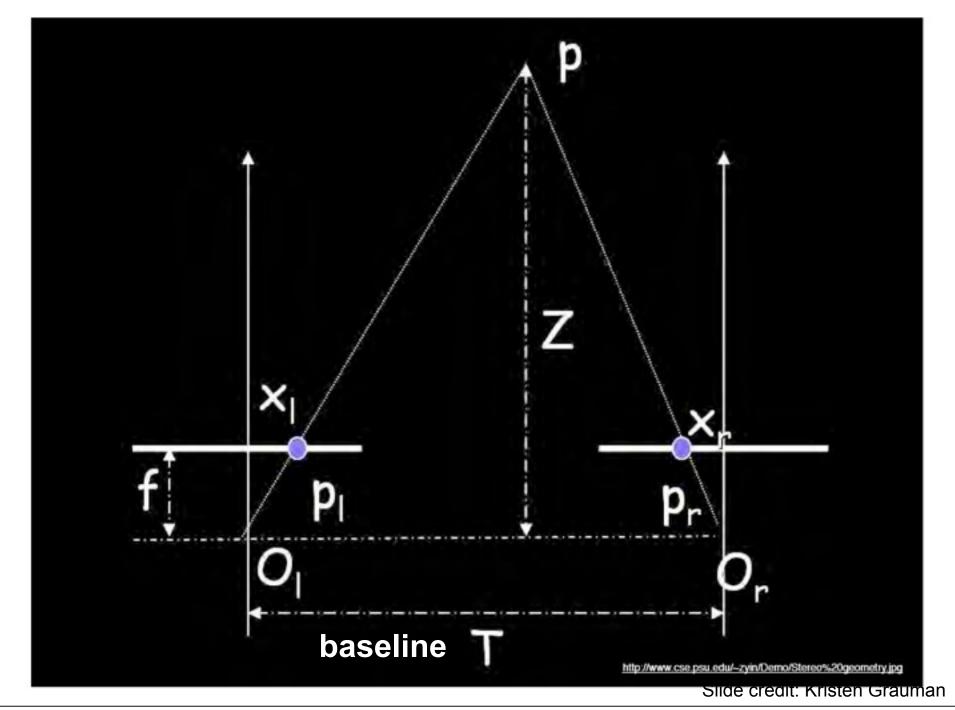
Stereo Topics

- Special, simple system, main idea
- More general camera conditions, epipolar constraints
 - epipolar geometry
 - epipolar algebra
- Image rectification
- Stereo matching (likelihood term)
- Stereo regularization (prior term)
- Inference
 - dynamic programming
 - graph cuts
- Structured light

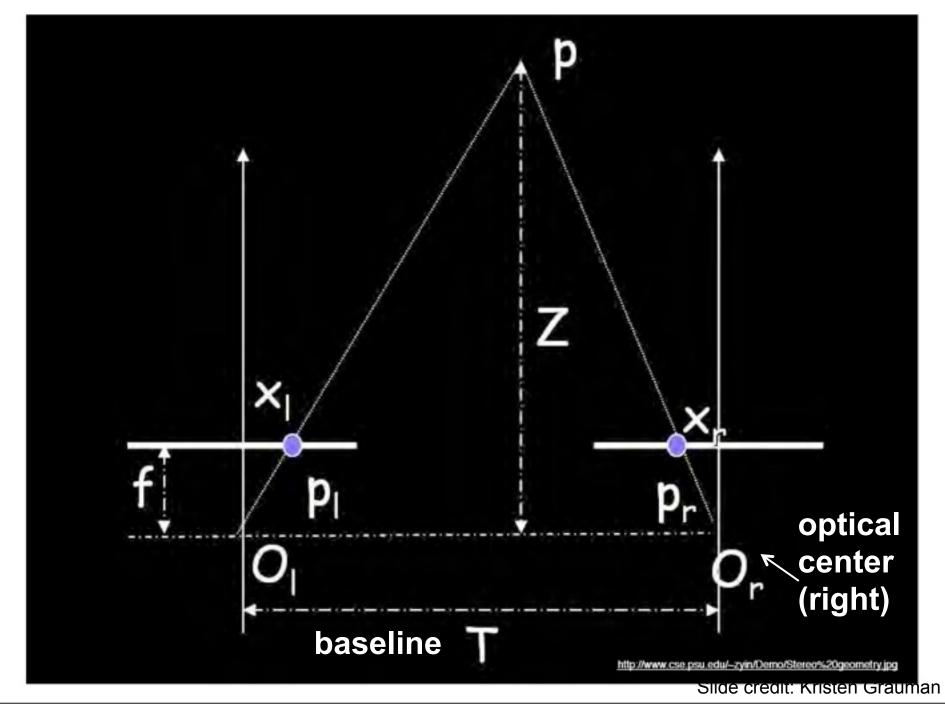
• First, assuming parallel optical axes, known camera parameters (i.e., calibrated cameras):



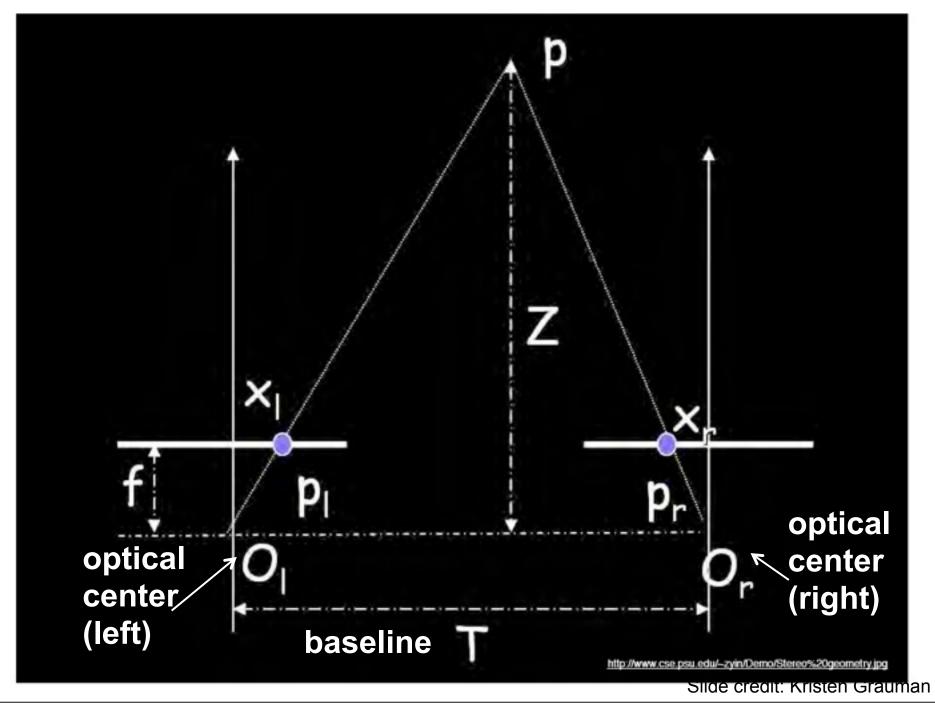
Monday, March 14, 2011

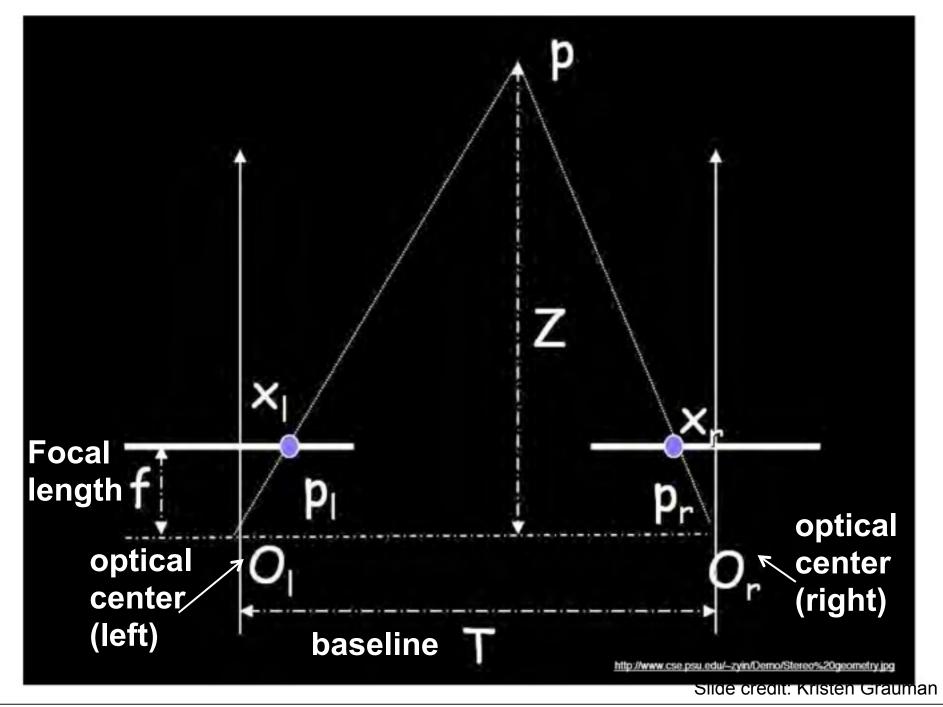


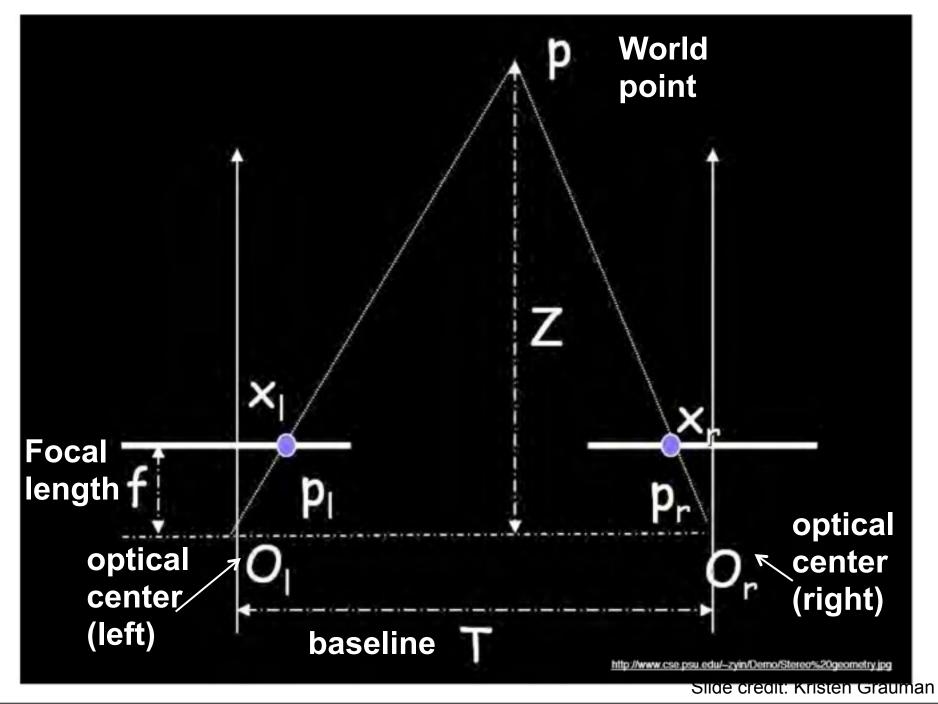
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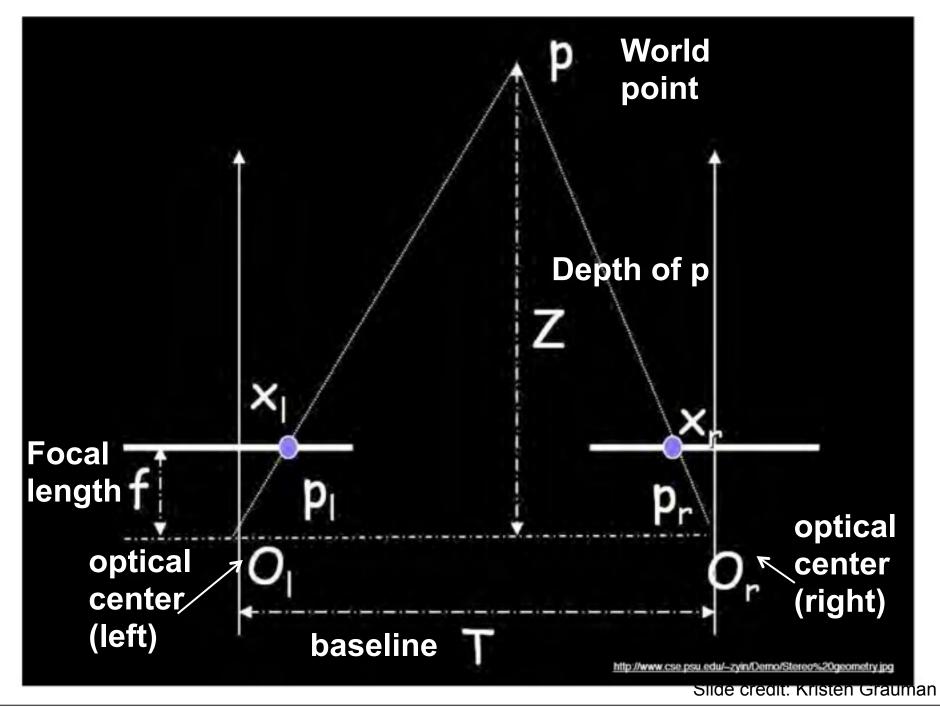


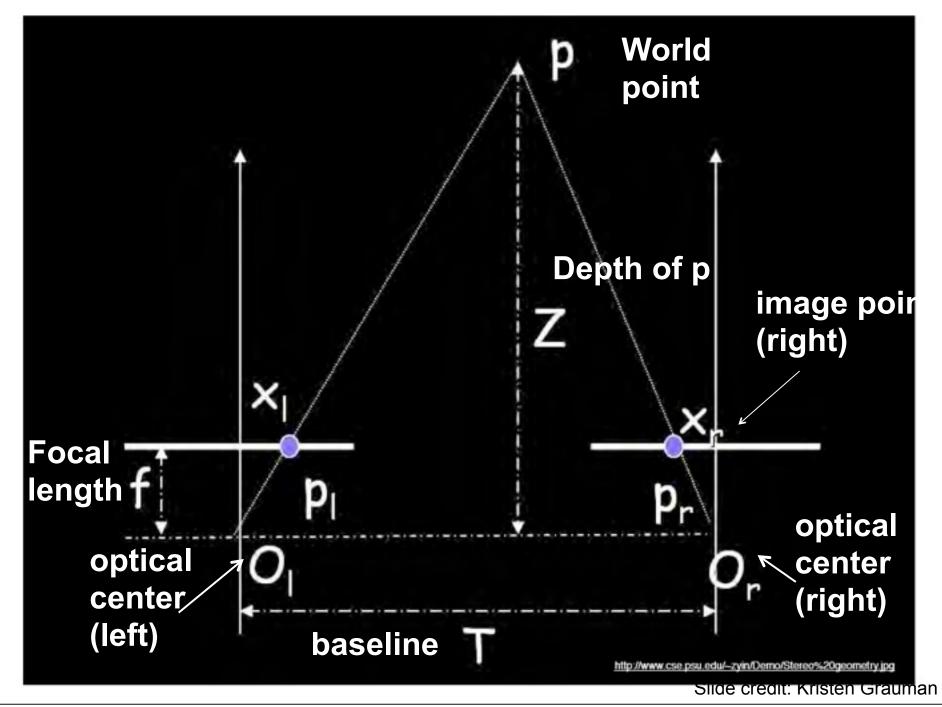
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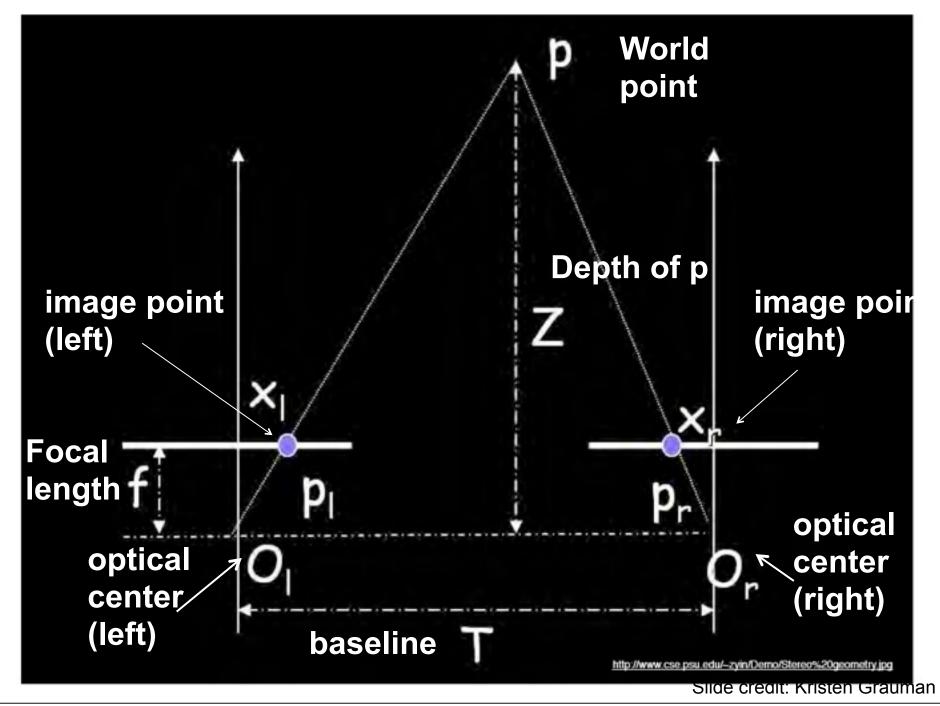


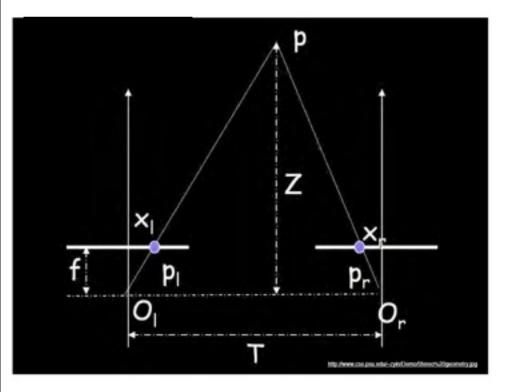


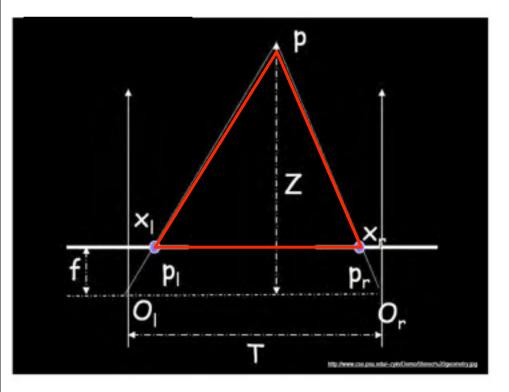


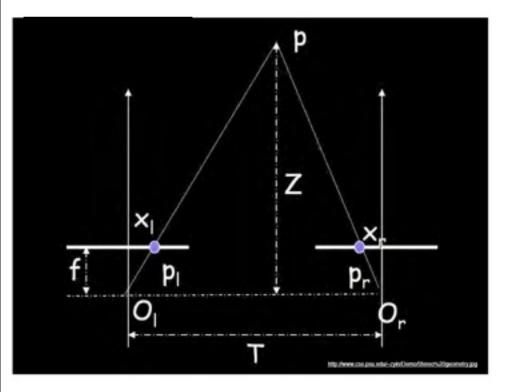


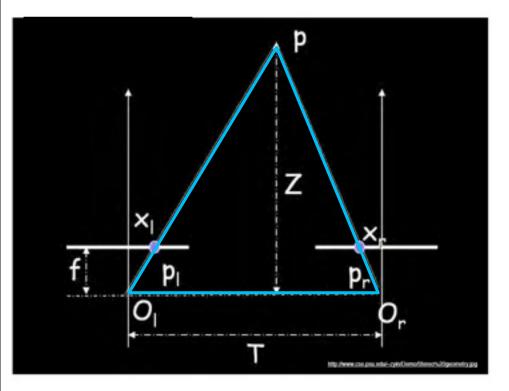


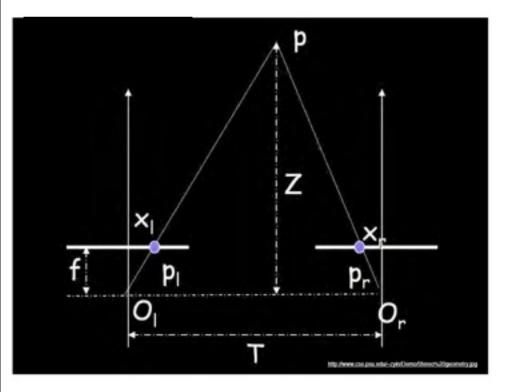




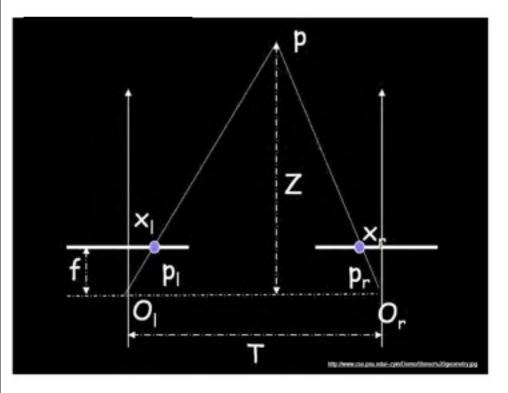






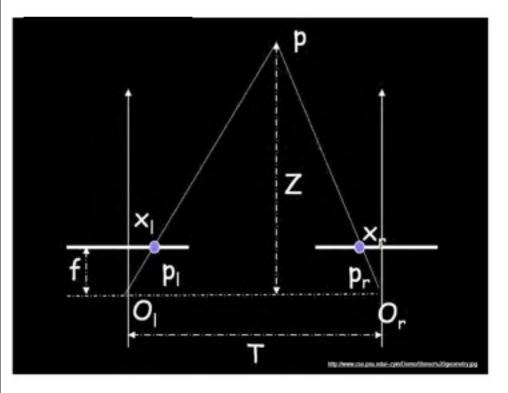


 Assume parallel optical axes, known camera parameters (i.e., calibrated cameras). We can triangulate via:



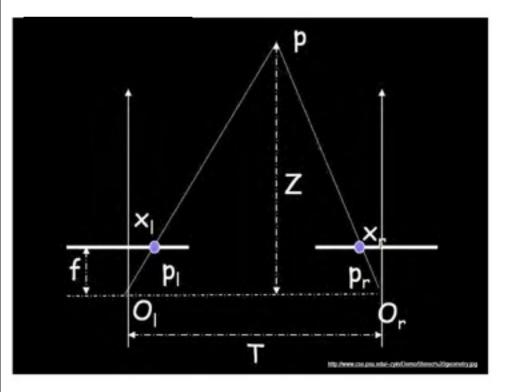
Similar triangles (p_1 , P, p_r) and (O_1 , P, O_r):

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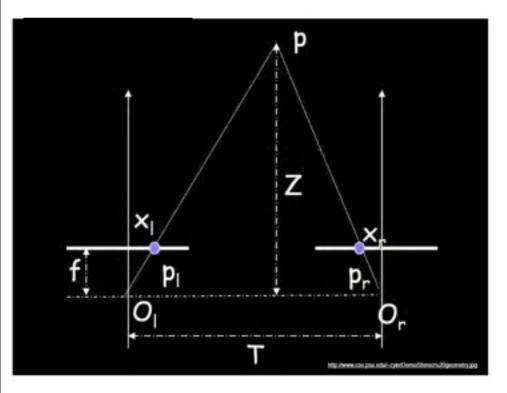
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Similar triangles (p_l, P, p_r) and (O_l, P, O_r) :

$$\frac{T + x_l - x_r}{Z - f} = \frac{T}{Z}$$

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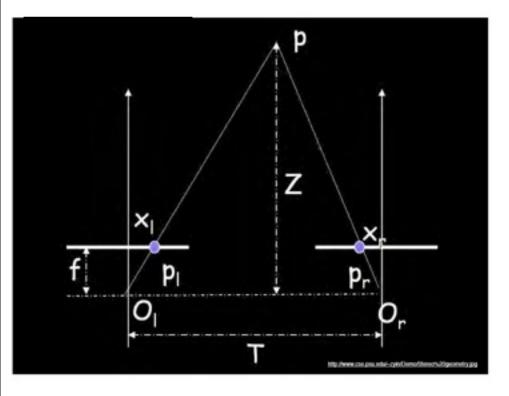


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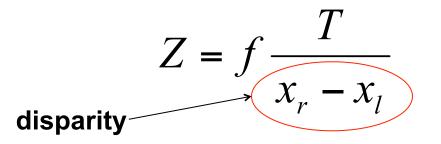
$$Z = f \frac{T}{x_r - x_l}$$

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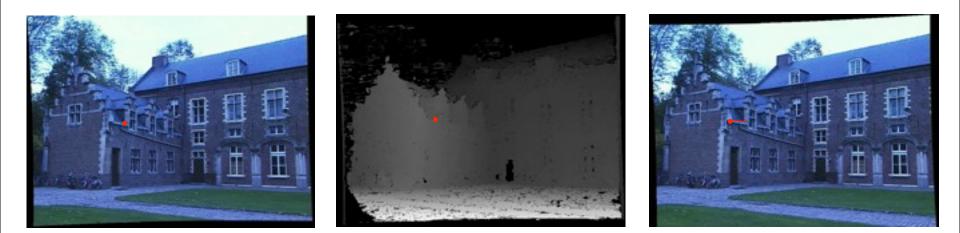


Depth from disparity

image I(x,y)

Disparity map D(x,y)

image l´(x´,y´)



(x´,y`)=(x+D(x,y), y)

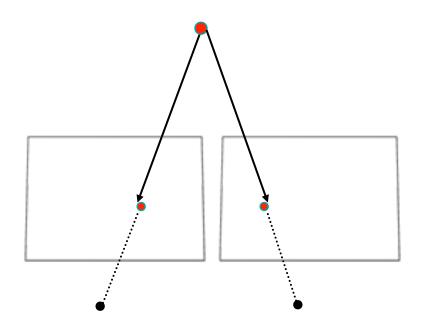
Slide credit: Kristen Grauman

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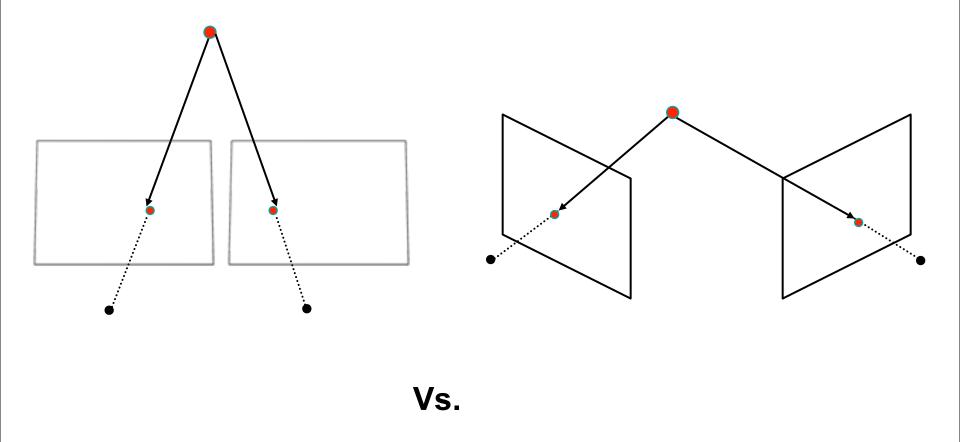
General case, with calibrated cameras

• The two cameras need not have parallel optical axes.

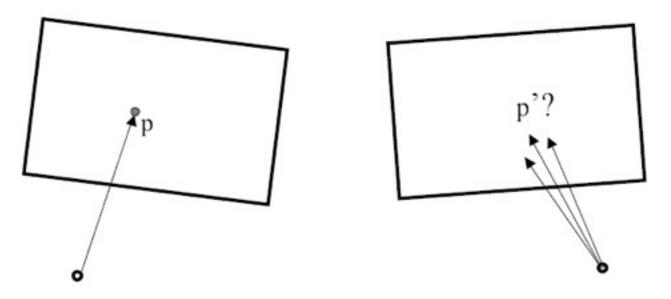


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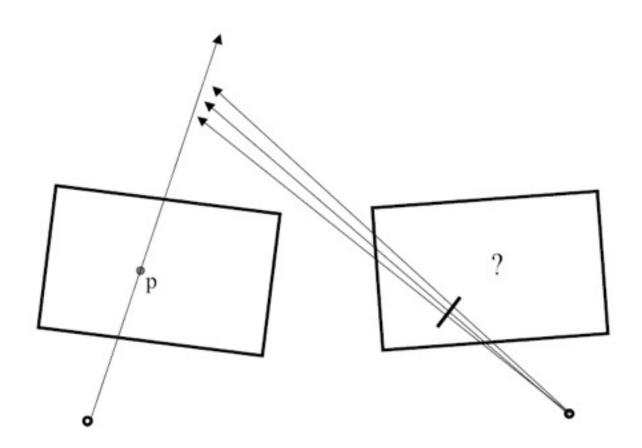


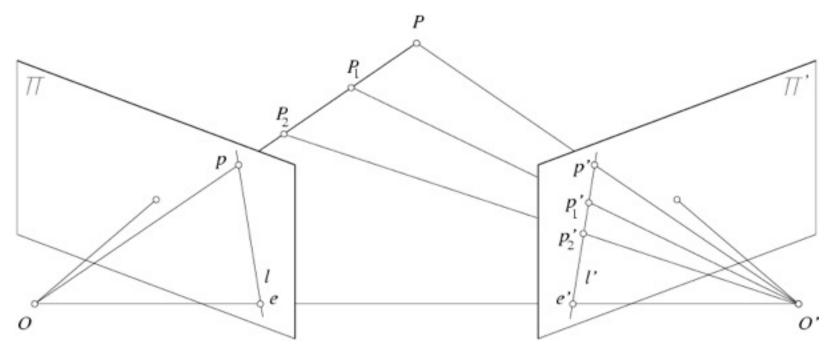
Stereo correspondence constraints

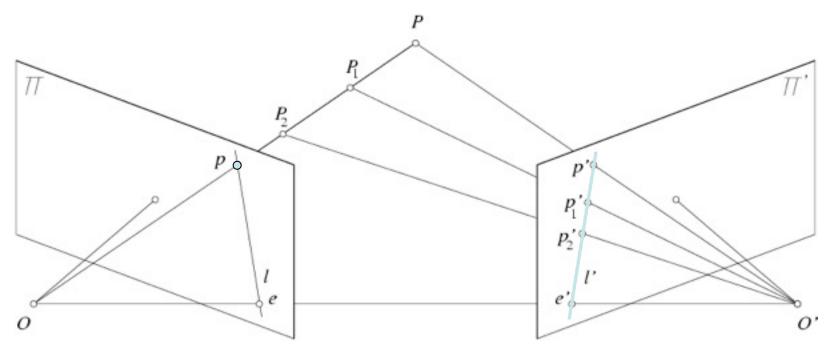


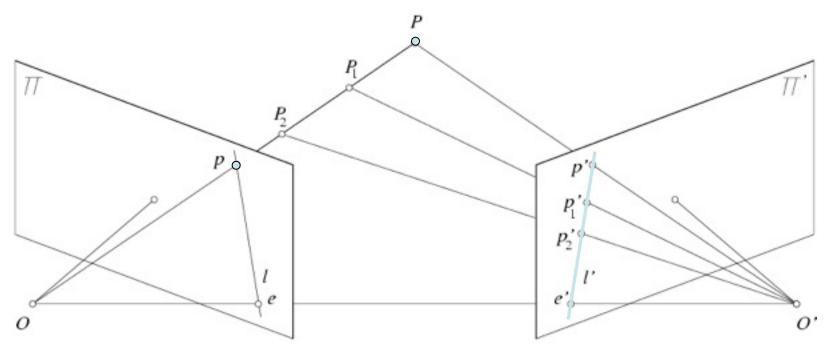
• Given p in left image, where can corresponding point p' be?

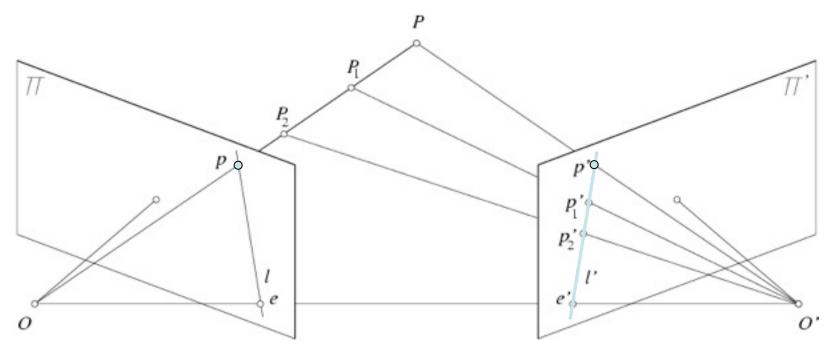
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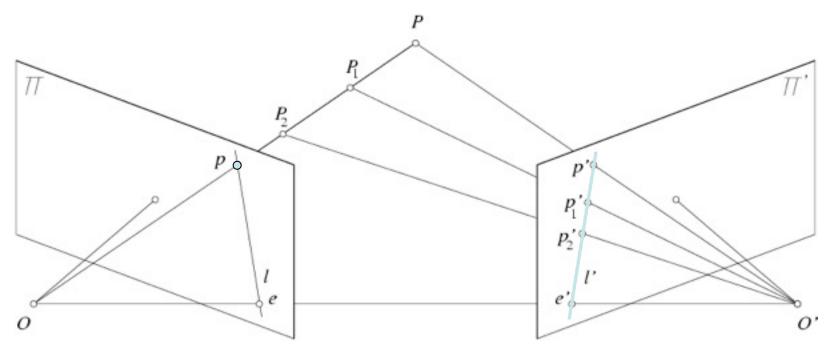


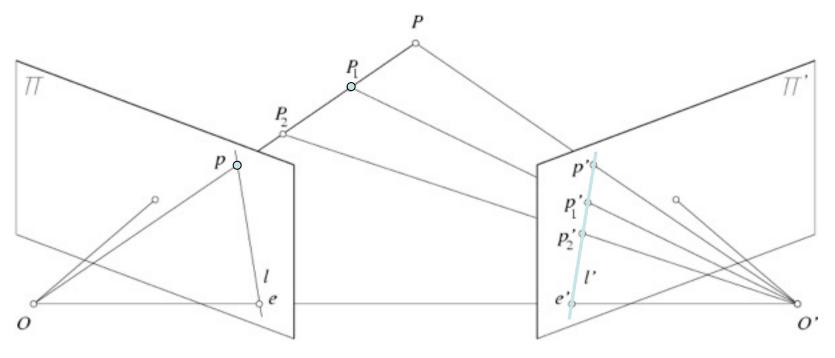


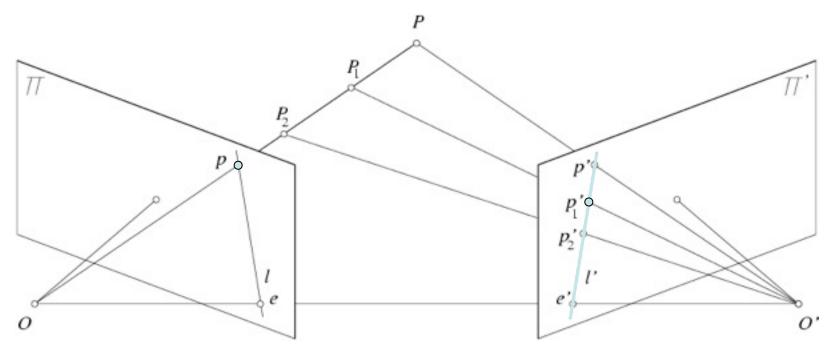


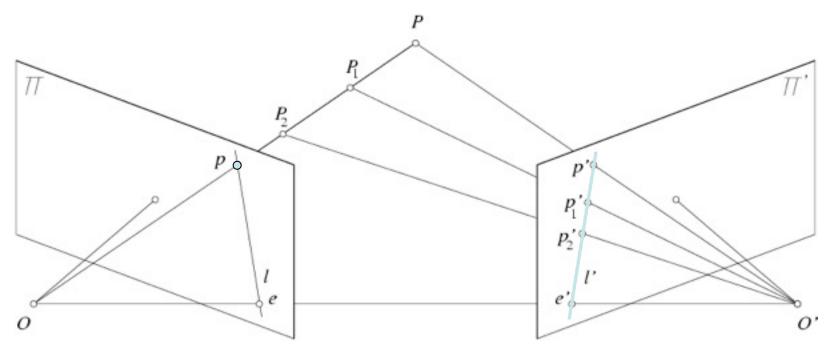


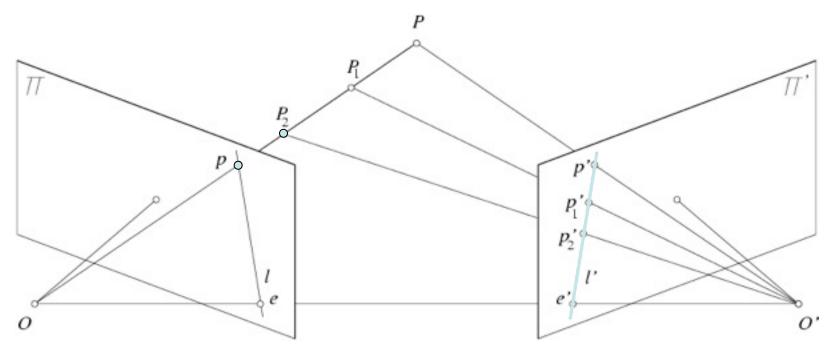


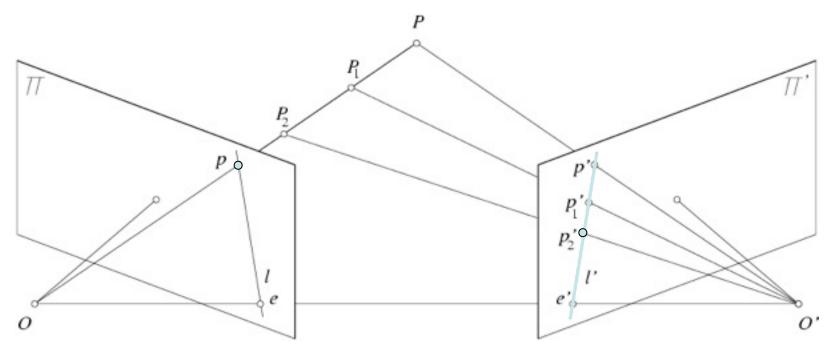


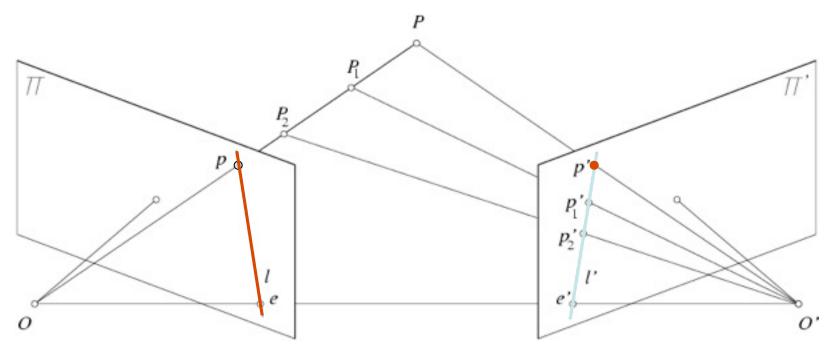


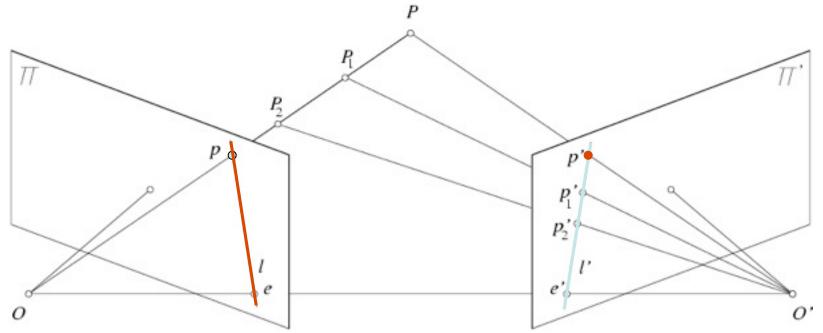




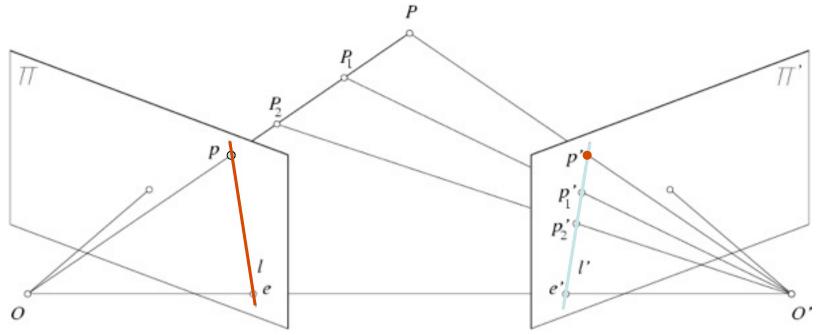






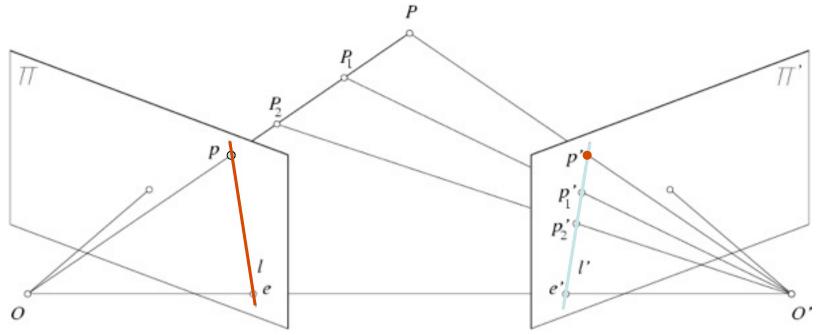


Geometry of two views constrains where the corresponding pixel for some image point in the first view must occur in the second view:



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• It must be on the line carved out by a plane connecting the world point and optical centers.



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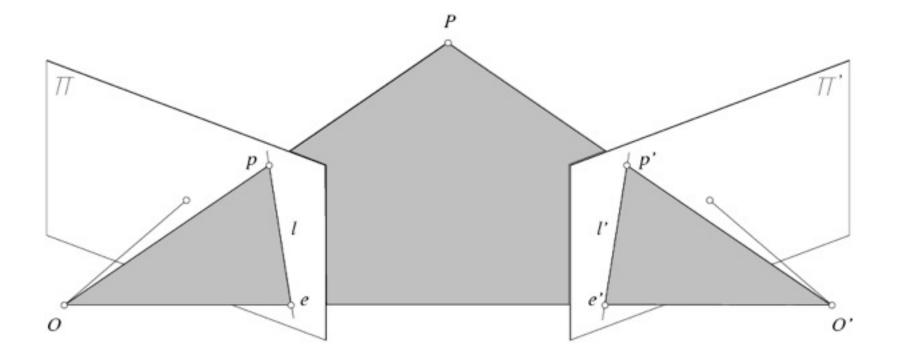
Why is this useful?



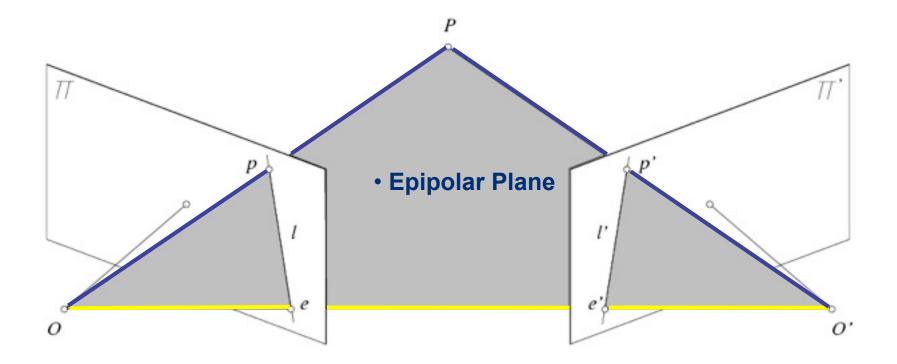
This is useful because it reduces the correspondence problem to a 1D search along an epipolar line.

Image from Andrew Zisserman

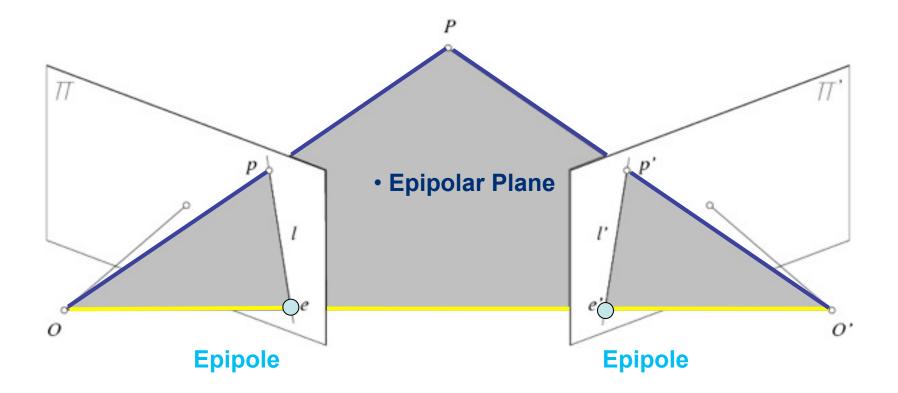
Slide credit: Kristen Grauman



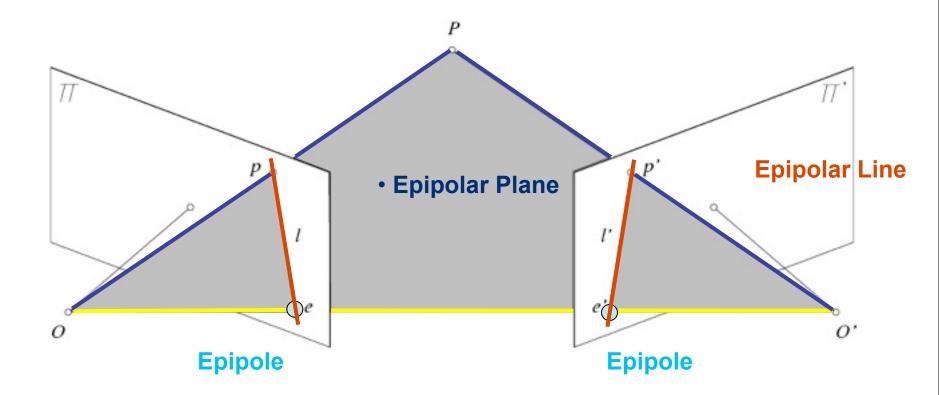
http://www.ai.sri.com/~luong/research/Meta3DViewer/EpipolarGeo.html



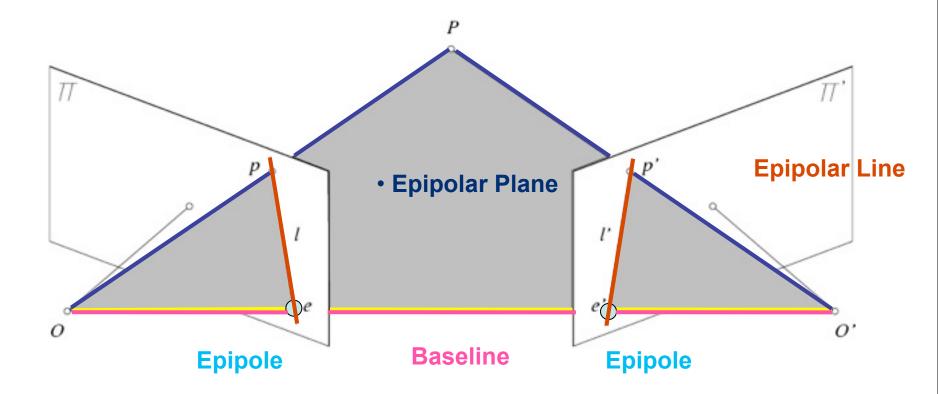
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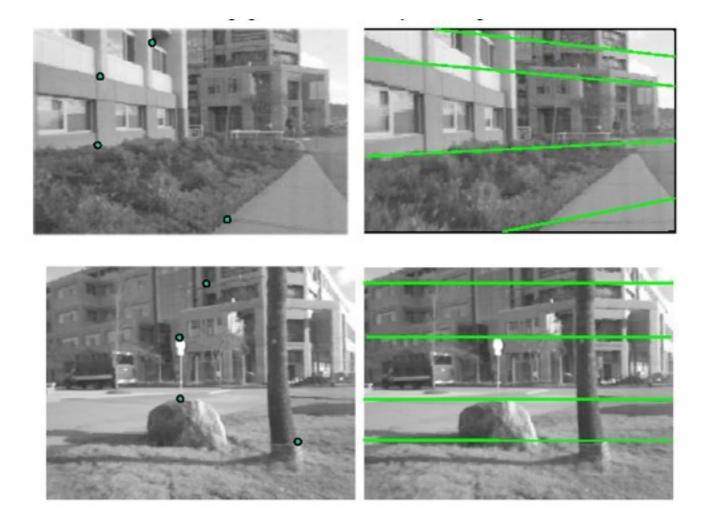


http://www.ai.sri.com/~luong/research/Meta3DViewer/EpipolarGeo.html

Epipolar geometry: terms

- **Baseline**: line joining the camera centers
- **Epipole**: point of intersection of baseline with the image plane
- Epipolar plane: plane containing baseline and world point
- Epipolar line: intersection of epipolar plane with the image plane
- All epipolar lines intersect at the epipole
- An epipolar plane intersects the left and right image planes in epipolar lines

Example



Example: converging cameras

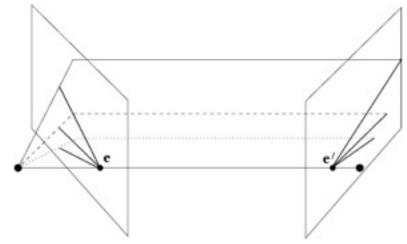


Figure from Hartley & Zisserman

Slide credit: Kristen Grauman

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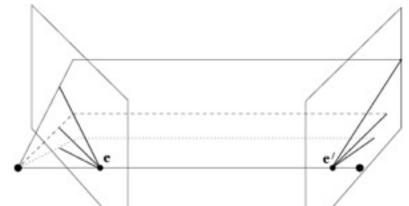


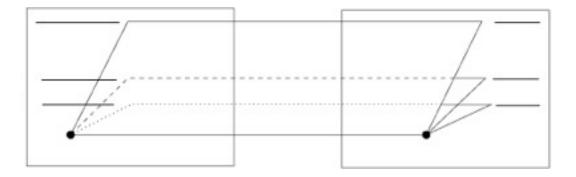


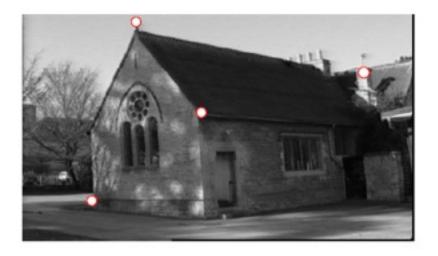


Figure from Hartley & Zisserman

Slide credit: Kristen Grauman

Example: parallel cameras





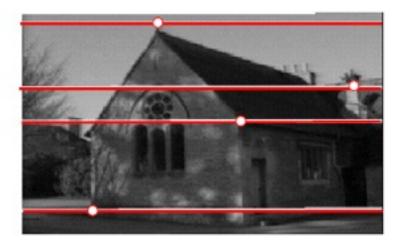
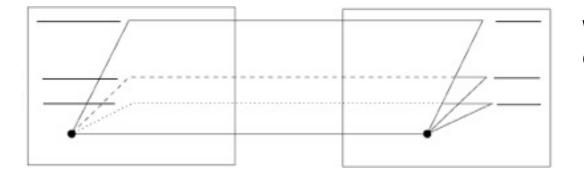


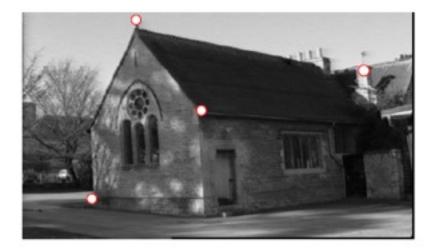
Figure from Hartley & Zisserman

Slide credit: Kristen Grauman

Example: parallel cameras



Where are the epipoles?



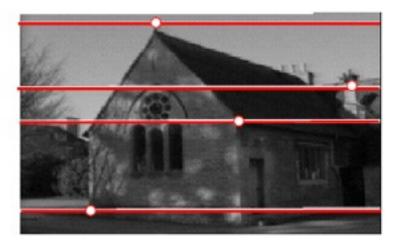


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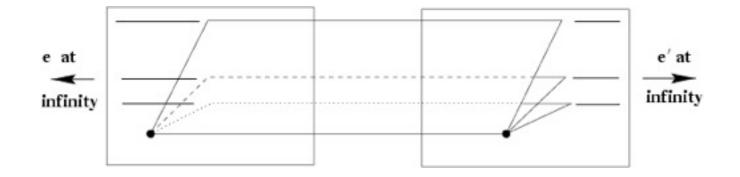




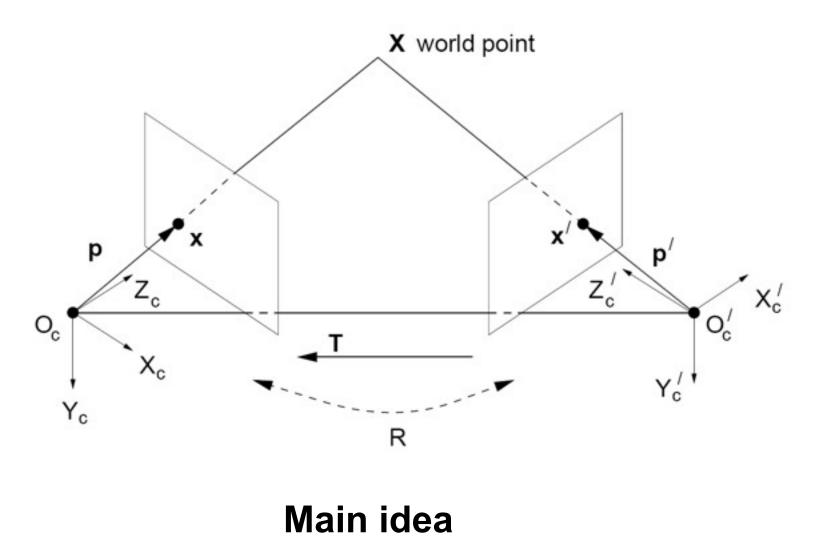
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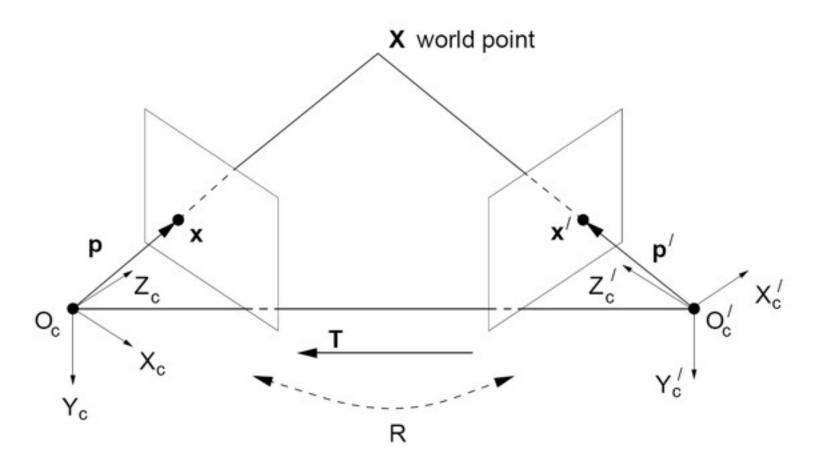
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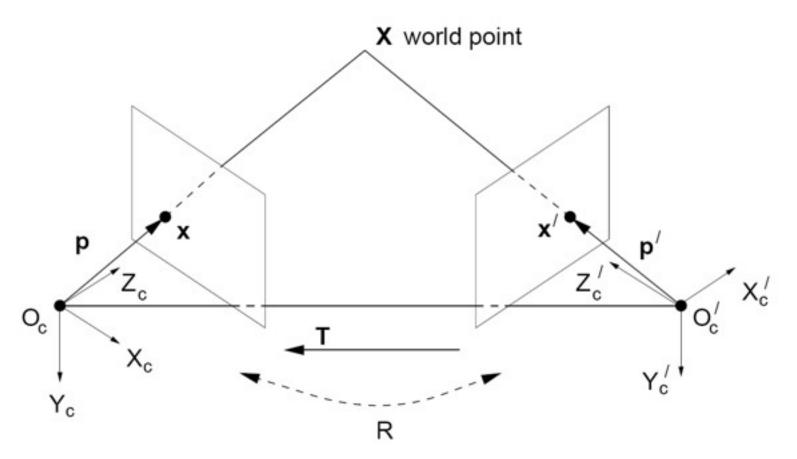
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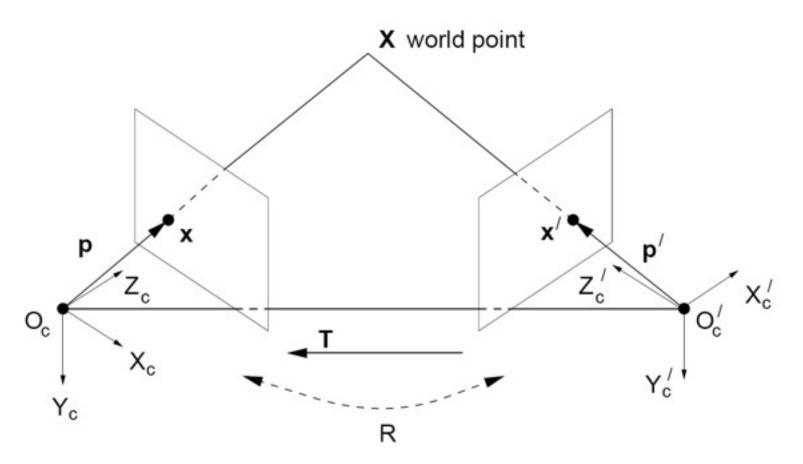
- So far, we have the explanation in terms of geometry.
- Now, how to express the epipolar constraints algebraically?



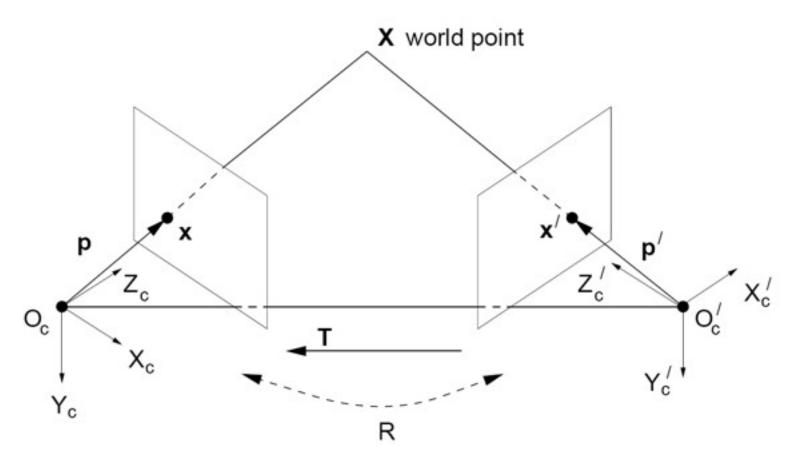




If the stereo rig is calibrated, we know : how to rotate and translate camera reference frame 1 to get to camera reference frame 2.

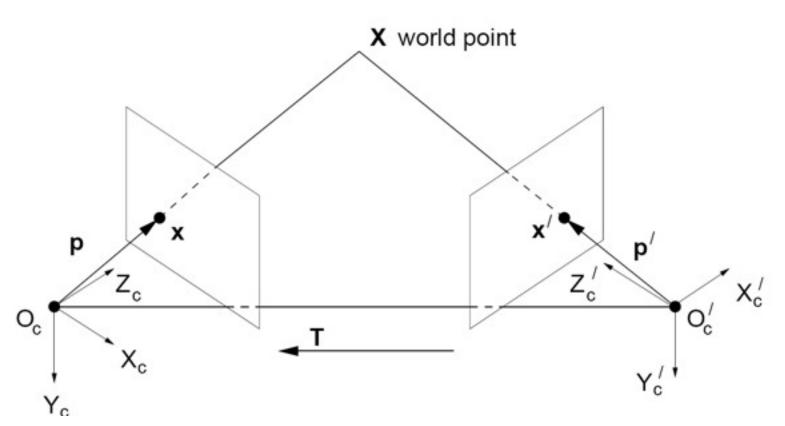


If the stereo rig is calibrated, we know : how to rotate and translate camera reference frame 1 to get to camera reference frame 2. Rotation: 3 x 3 matrix R; translation: 3 vector T.



If the stereo rig is calibrated, we know : how to rotate and translate camera reference frame 1 to get to camera reference frame 2. $X'_{c} = RX_{c} + T'$

From geometry to algebra

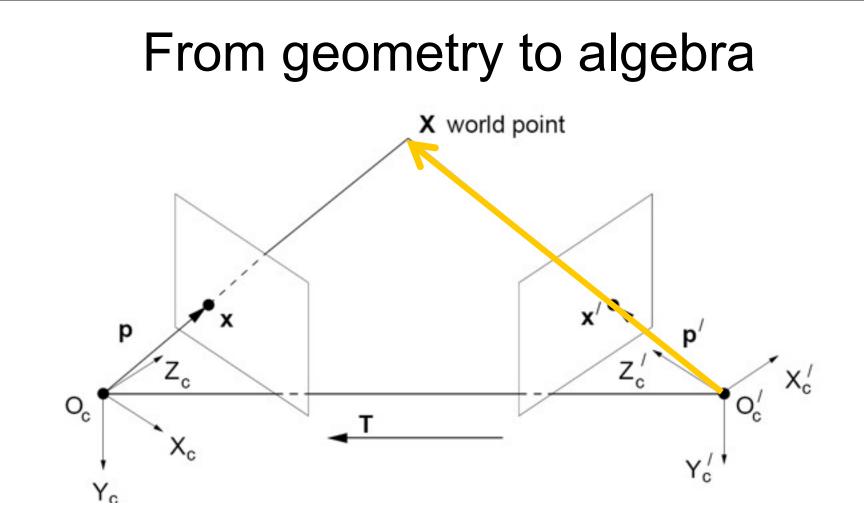


$$\begin{aligned} X' &= RX + T' \\ T' \times X' &= T' \times RX + T' \times T' \\ &= T' \times RX \\ X' \cdot (T' \times X') &= X' \cdot (T' \times RX) = 0 \end{aligned}$$

From unprimed to primed coordinate system

Cross with T' on both sides

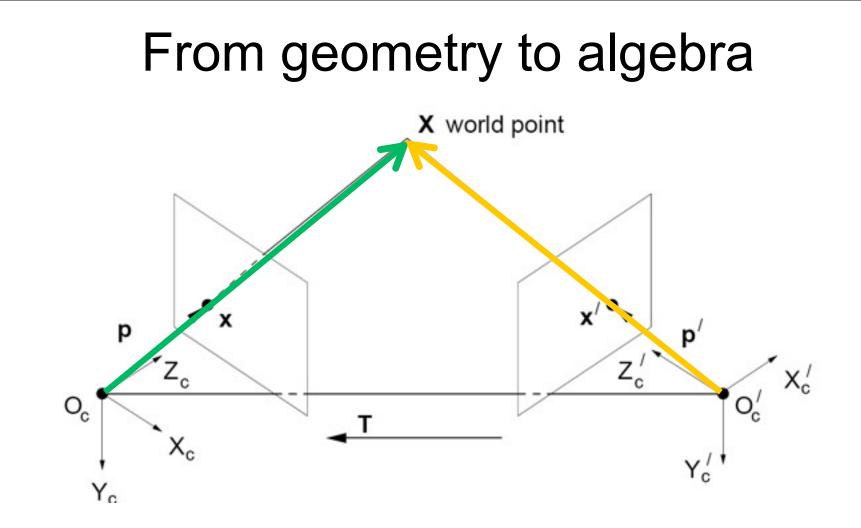
Simplify



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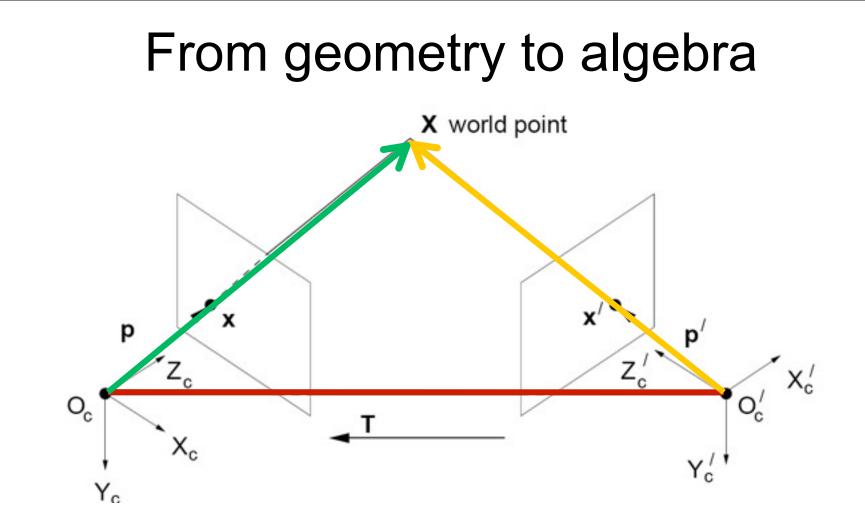
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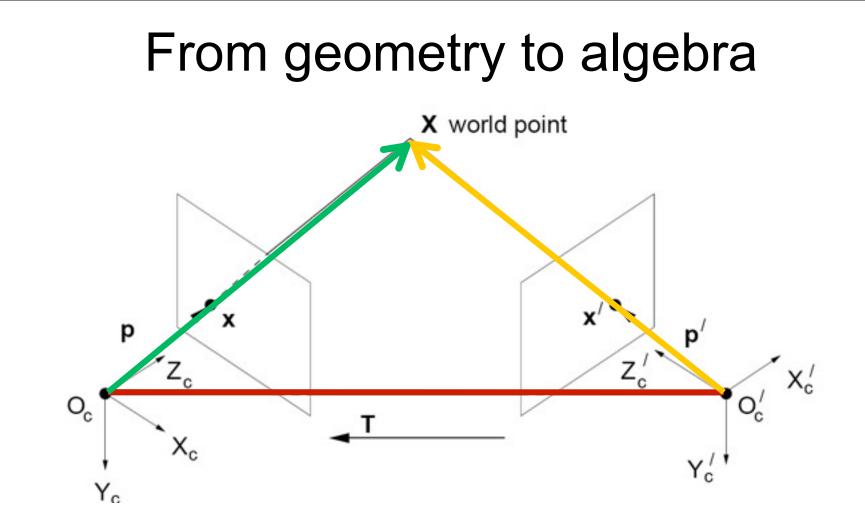
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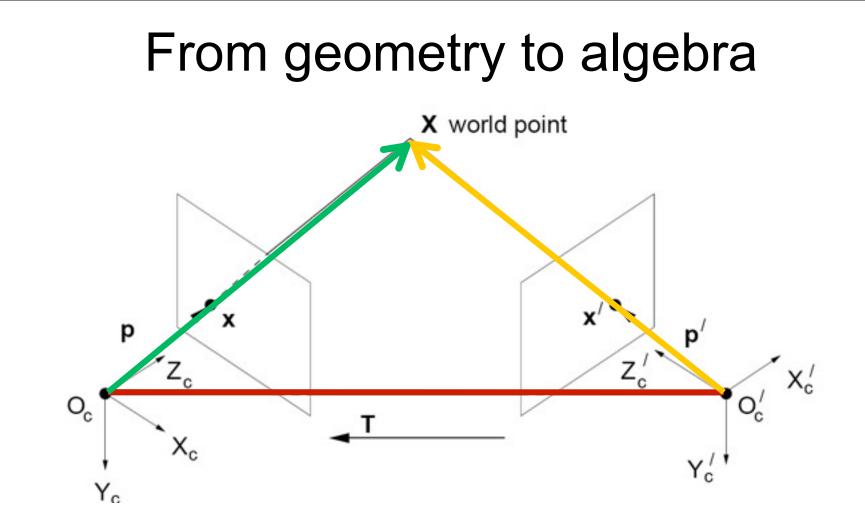
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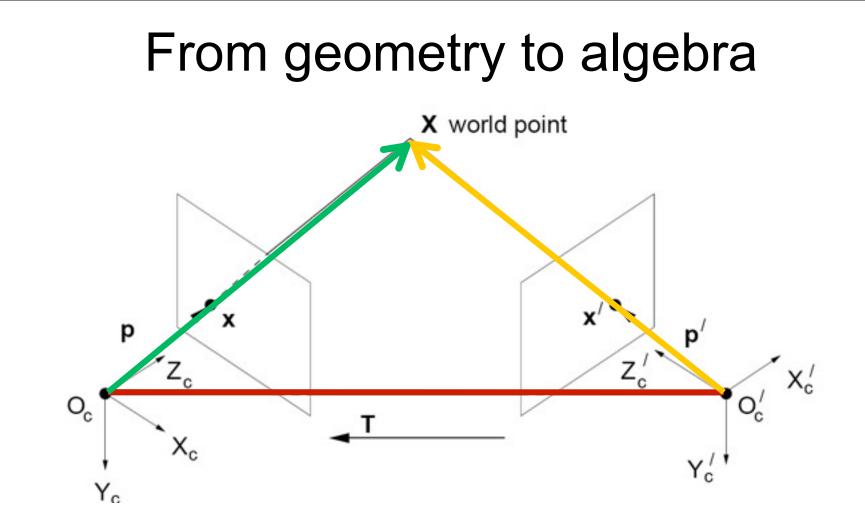
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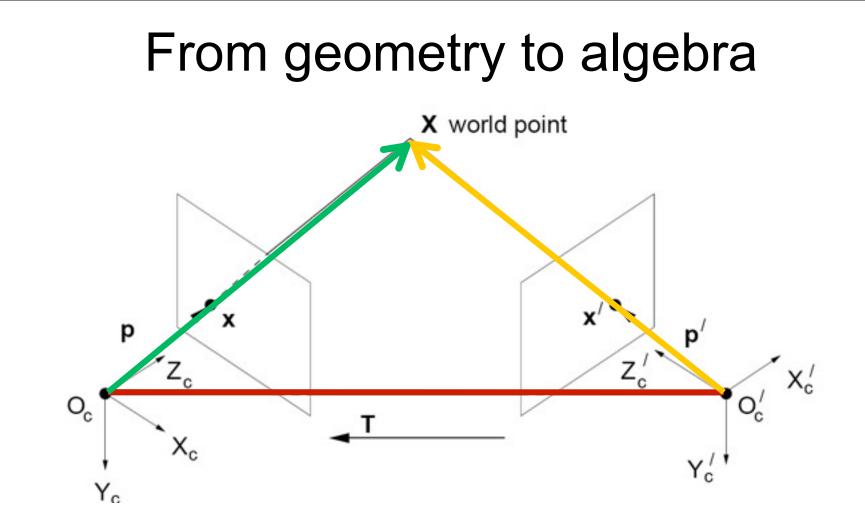
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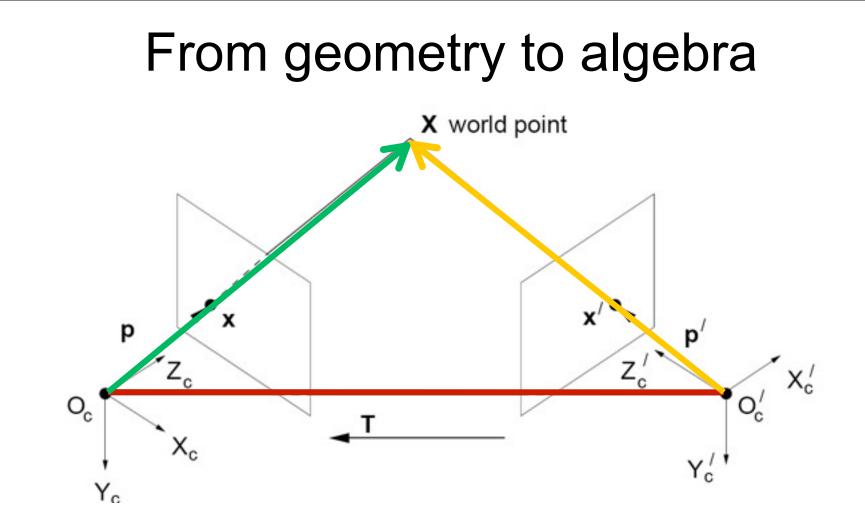
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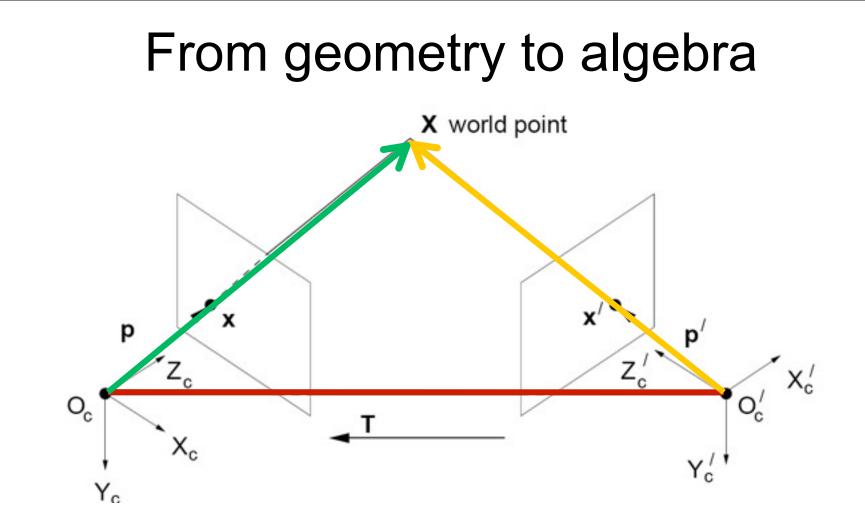
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Cross with T' on both sides

Simplify

Aside: cross product

$\vec{a} \times \vec{b} = \vec{c}$

Aside: cross product

$$\vec{a} \times \vec{b} = \vec{c} \qquad \qquad \vec{a} \cdot \vec{c} = 0 \\ \vec{b} \cdot \vec{c} = 0$$

Vector cross product takes two vectors and returns a third vector that's perpendicular to both inputs.

So here, c is perpendicular to both a and b, which means the dot product = 0.

$$\vec{a} \times \vec{b} = \begin{bmatrix} 0 & -a_3 & a_2 \\ a_3 & 0 & -a_1 \\ -a_2 & a_1 & 0 \end{bmatrix} \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix} = \vec{c} \qquad \vec{a} \cdot \vec{c} = \vec{b} \cdot \vec{c} = \vec{c}$$

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0

O

$$\vec{a} \times \vec{b} = \begin{bmatrix} 0 & -a_3 & a_2 \\ a_3 & 0 & -a_1 \\ -a_2 & a_1 & 0 \end{bmatrix} \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix} = \vec{c} \qquad \vec{a} \cdot \vec{c} = \mathbf{0}$$

Can be expressed as a matrix multiplication.

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Can be expressed as a matrix multiplication.

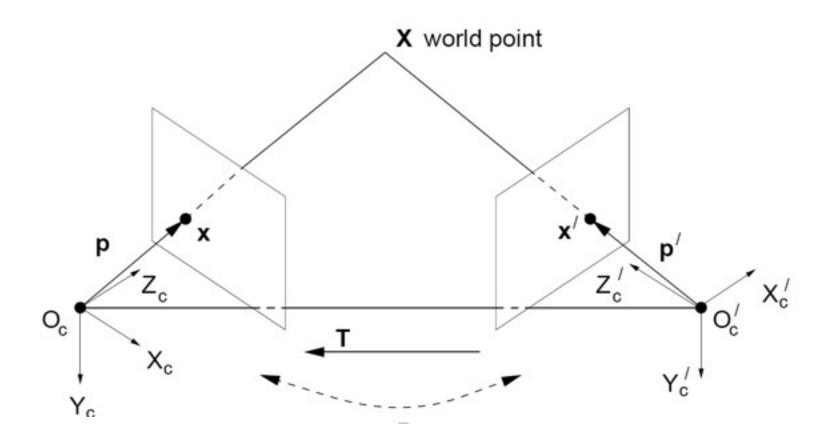
$$\begin{bmatrix} a_x \end{bmatrix} = \begin{bmatrix} 0 & -a_3 & a_2 \\ a_3 & 0 & -a_1 \\ -a_2 & a_1 & 0 \end{bmatrix}$$

$$\vec{a} \times \vec{b} = \begin{bmatrix} 0 & -a_3 & a_2 \\ a_3 & 0 & -a_1 \\ -a_2 & a_1 & 0 \end{bmatrix} \begin{bmatrix} b_1 \\ b_2 \\ b_3 \end{bmatrix} = \vec{c} \qquad \vec{a} \cdot \vec{c} = \mathbf{0}$$

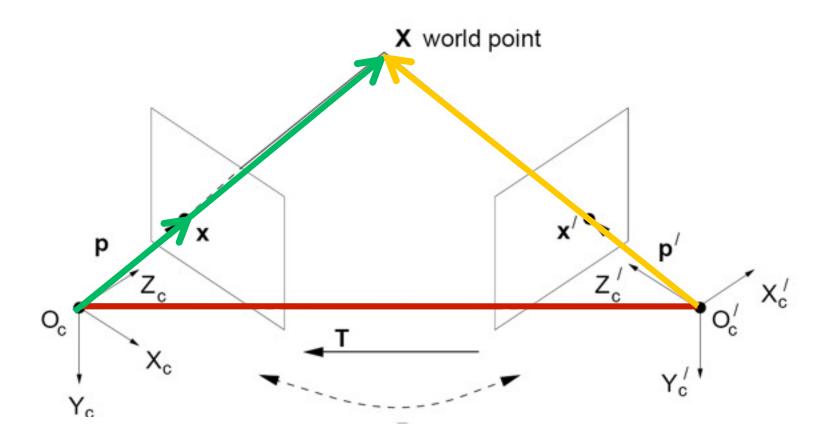
Can be expressed as a matrix multiplication.

$$\begin{bmatrix} a_x \end{bmatrix} = \begin{bmatrix} 0 & -a_3 & a_2 \\ a_3 & 0 & -a_1 \\ -a_2 & a_1 & 0 \end{bmatrix} \qquad \begin{bmatrix} \vec{a} \times \vec{b} = [a_x] \vec{b} \end{bmatrix}$$

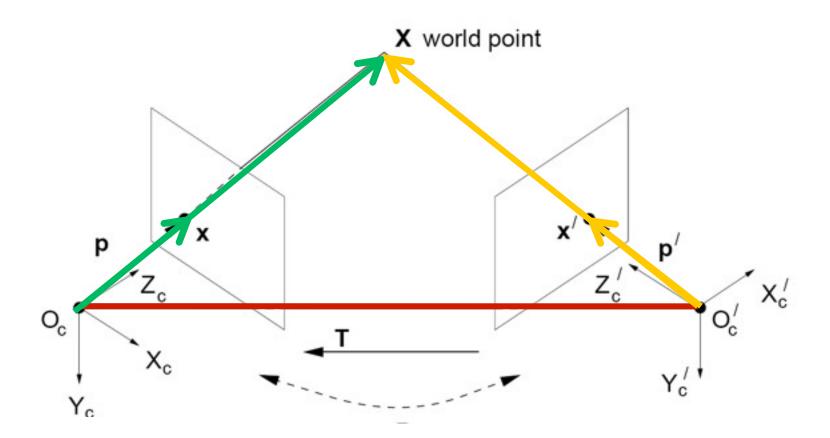
x and x' are scaled versions of X and X'

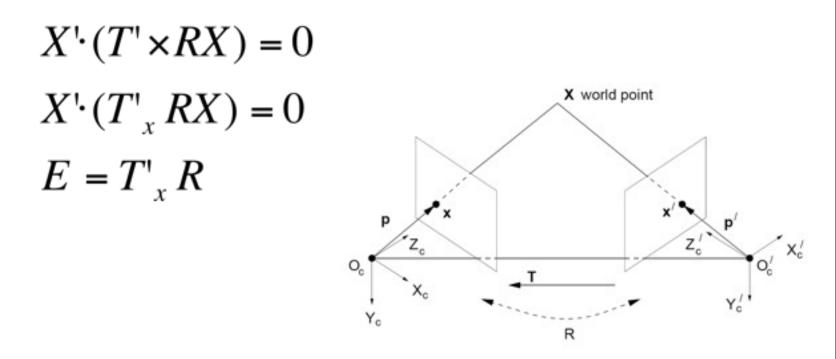


x and x' are scaled versions of X and X'

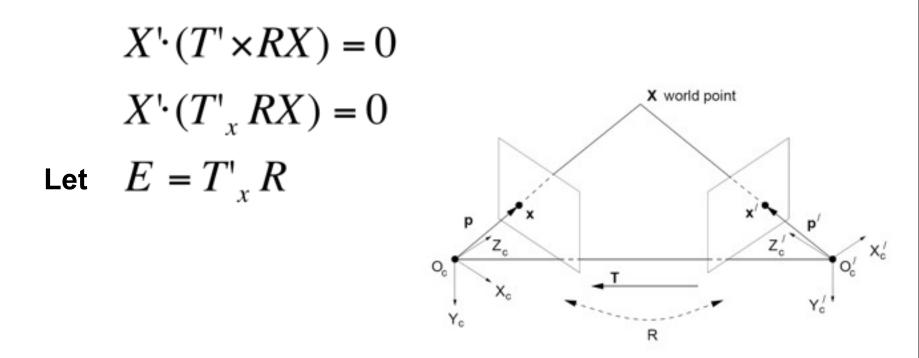


x and x' are scaled versions of X and X'

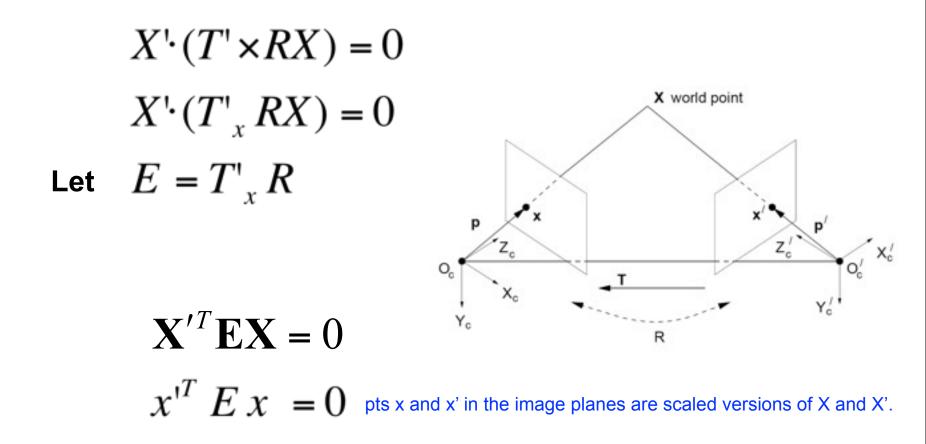




 $x'^T E x = 0$ pts x and x' in the image planes are scaled versions of X and X'.



 $x'^T E x = 0$ pts x and x' in the image planes are scaled versions of X and X'.



E is called the essential matrix, and it relates corresponding image points between both cameras, given the rotation and translation.

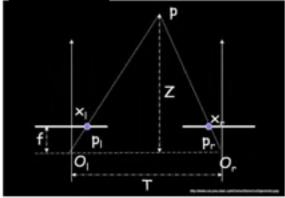
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If we observe a point in one image, its position in the other image is constrained to lie on line defined by above.

E is called the essential matrix, and it relates corresponding image points between both cameras, given the rotation and translation.

If we observe a point in one image, its position in the other image is constrained to lie on line defined by above.

Note: these points are in camera coordinate systems.



R =T = $E = [T_x]R =$

 $\mathbf{E} = [\mathbf{T}_x]\mathbf{R} =$

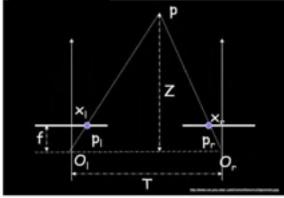
 $\mathbf{T} =$

Z

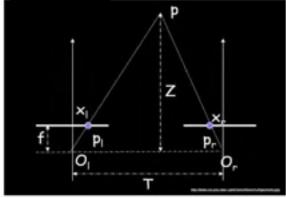
т

Pi

0



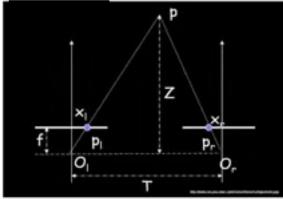
 $\mathbf{R} = \mathbf{I}$ $\mathbf{T} = [-d, 0, 0]^{\mathrm{T}}$ $\mathbf{E} = [\mathbf{T}_{\mathrm{x}}]\mathbf{R} =$



$$R = I$$

$$T = [-d,0,0]^{T}$$

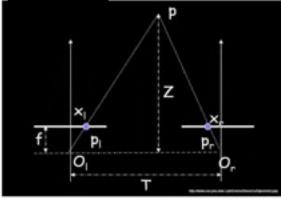
$$E = [T_{x}]R = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix}$$



$$R = I \qquad p = [x, y, f]$$

$$T = [-d, 0, 0]^{T} \qquad p' = [x', y', f]$$

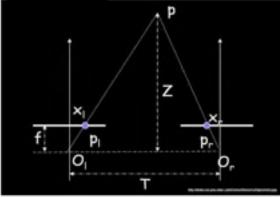
$$E = [T_x]R = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix}$$



Calleras
R = **I**
p = [x, y, f]
T =
$$[-d, 0, 0]^{T}$$

p' = [x', y', f]
E = [**T**_x]**R** = $\begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{pmatrix}$

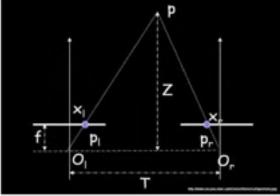
 $\mathbf{p}^{'^{\mathrm{T}}}\mathbf{E}\mathbf{p}=\mathbf{0}$



R = **I**
T =
$$[-d,0,0]^{T}$$

E = $[\mathbf{T}_{x}]\mathbf{R} = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix}$
p = $[x, y, f]$
p' = $[x', y', f]$

$$\mathbf{p}^{T}\mathbf{E}\mathbf{p} = \mathbf{0} \qquad \begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ f \end{bmatrix} = \mathbf{0}$$



$$R = I p = [x, y, f]$$

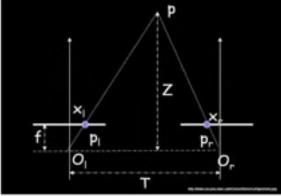
$$T = [-d, 0, 0]^{T} p' = [x', y', f]$$

$$E = [T_x]R = \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix}$$

$$\mathbf{p}^{T}\mathbf{E}\mathbf{p} = \mathbf{0}$$

$$\begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ f \end{bmatrix} = \mathbf{0}$$

$$\Leftrightarrow \begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 \\ df \\ -dy \end{bmatrix} = \mathbf{0}$$



$$R = I p = [x, y, f]$$

$$T = [-d, 0, 0]^{T} p' = [x', y', f]$$

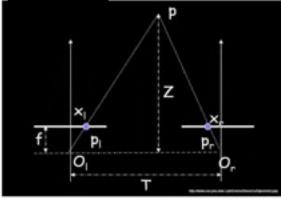
$$E = [T_x]R = \begin{pmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{pmatrix}$$

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$$\begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ f \end{bmatrix} = \mathbf{0}$$

$$\Leftrightarrow \begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 \\ df \\ -dy \end{bmatrix} = \mathbf{0}$$

$$\Leftrightarrow y = y'$$



$$\mathbf{R} = \mathbf{I}$$
$$\mathbf{T} = \begin{bmatrix} -d, 0, 0 \end{bmatrix}^{\mathrm{T}}$$
$$\mathbf{E} = \begin{bmatrix} \mathbf{T}_{\mathbf{x}} \end{bmatrix} \mathbf{R} = \begin{bmatrix} \mathbf{0} & \mathbf{0} & \mathbf{0} \\ \mathbf{0} & \mathbf{0} & \mathbf{d} \\ \mathbf{0} - \mathbf{d} & \mathbf{0} \end{bmatrix}$$

$$\mathbf{p} = [x, y, f]$$

 $\mathbf{p'} = [x', y', f]$

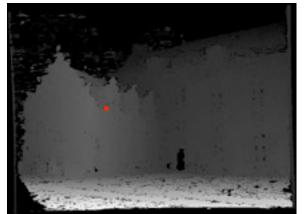
 $\mathbf{p}^{T}\mathbf{E}\mathbf{p} = 0$ $\begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 & 0 & 0 \\ 0 & 0 & d \\ 0 & -d & 0 \end{bmatrix} \begin{bmatrix} x \\ y \\ f \end{bmatrix} = 0$ For the parallel cameras, image of any point must lie on same horizontal line in each image plane. $\Leftrightarrow \begin{bmatrix} x' & y' & f \end{bmatrix} \begin{bmatrix} 0 \\ df \\ -dy \end{bmatrix} = 0$ $\Leftrightarrow y = y'$

image I(x,y)

Disparity map D(x,y)

image l´(x´,y´)







(x',y')=(x+D(x,y),y)

Slide credit: Kristen Grauman

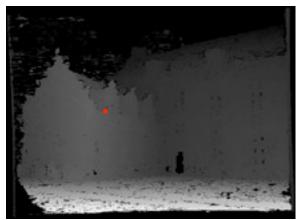
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image I(x,y)

Disparity map D(x,y)

image l´(x´,y´)







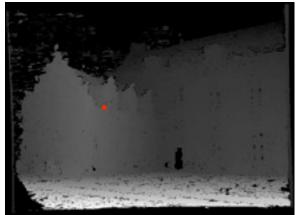
(x['],y['])=(x+D(x,y),y)

image I(x,y)

Disparity map D(x,y)

image l´(x´,y´)





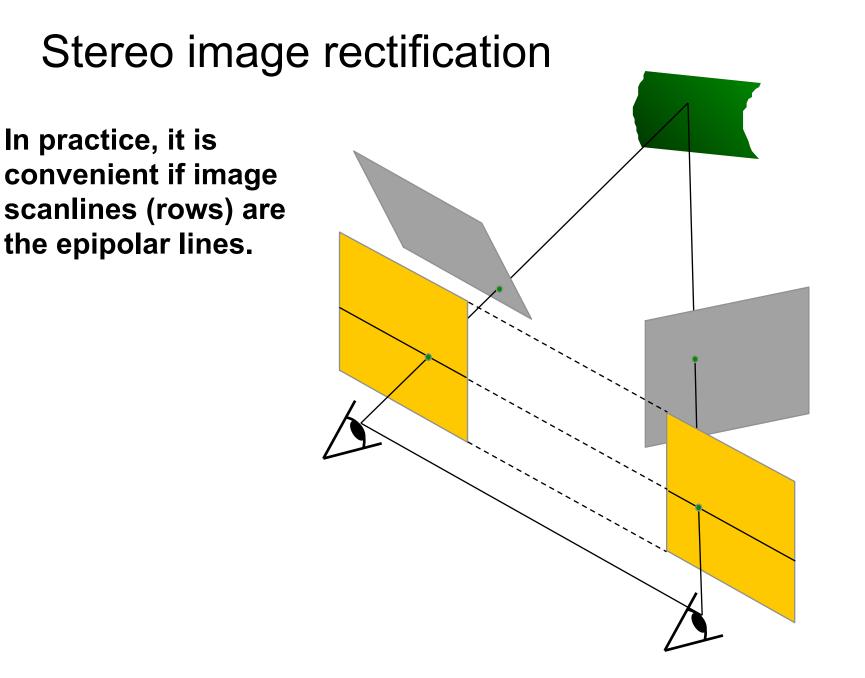


(x['],y['])=(x+D(x,y),y)

What about when cameras' optical axes are not parallel?

Stereo Topics

- Special, simple system, main idea
- More general camera conditions, epipolar constraints
 - epipolar geometry
 - epipolar algebra
- Image rectification
- Stereo matching (likelihood term)
- Stereo regularization (prior term)
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- Structured light



Adapted from Li Zhang

Slide credit: Kristen Grauman

Monday, March 14, 2011

Stereo image rectification

In practice, it is convenient if image scanlines (rows) are the epipolar lines.

Reproject image planes onto a common plane parallel to the line between optical

centers

Pixel motion is horizontal after this transformation

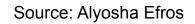
Two homographies (3x3 transforms), one for each input image reprojection

See Szeliski book, Sect. 2.1.5, Fig. 2.12, and "Mapping from one camera to another" p. 56

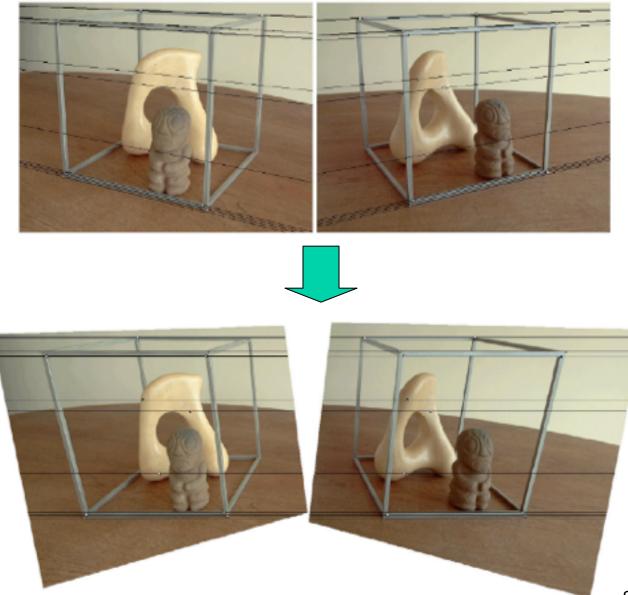
Adapted from Li Zhang

Stereo image rectification: example





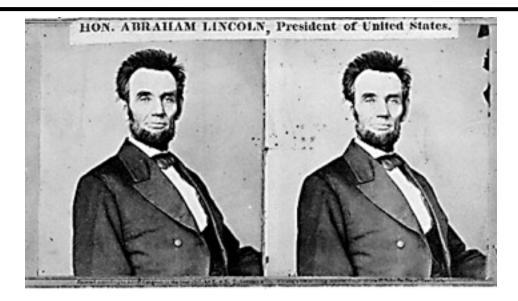
Stereo image rectification: example



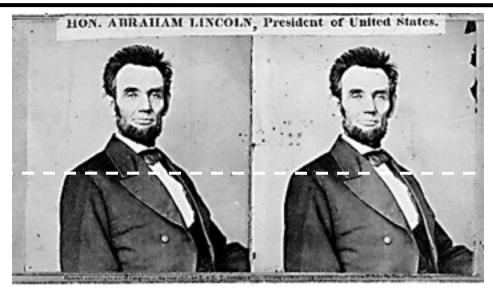
Source: Alyosha Efros

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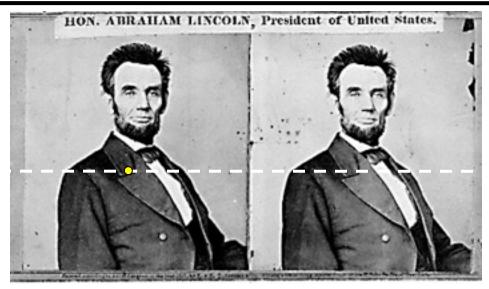


Slide credit: Rick Szeliski



For each epipolar line

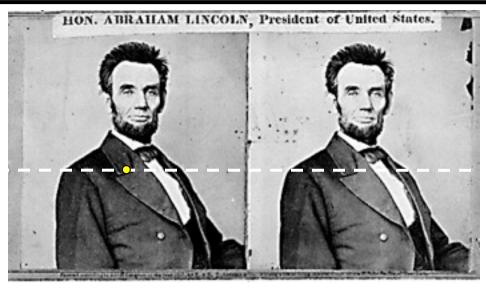
Slide credit: Rick Szeliski



For each epipolar line

For each pixel in the left image

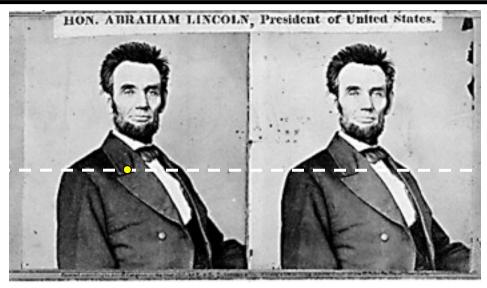
Slide credit: Rick Szeliski



For each epipolar line

For each pixel in the left image

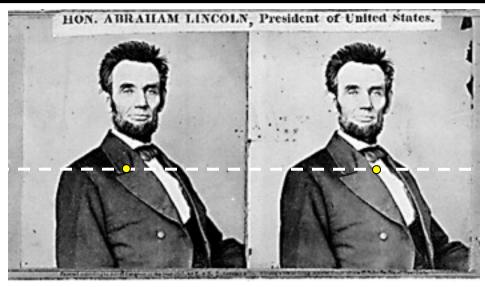
• compare with every pixel on same epipolar line in right image



For each epipolar line

For each pixel in the left image

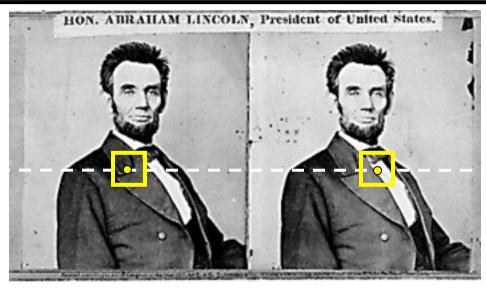
- compare with every pixel on same epipolar line in right image
- pick pixel with minimum match cost



For each epipolar line

For each pixel in the left image

- compare with every pixel on same epipolar line in right image
- pick pixel with minimum match cost



For each epipolar line

For each pixel in the left image

- compare with every pixel on same epipolar line in right image
- pick pixel with minimum match cost

Improvement: match windows

Slide credit: Rick Szeliski

Image block matching

How do we determine correspondences?

block matching or SSD (sum squared differences)

 $E(x, y; d) = \sum_{\substack{(x', y') \in N(x, y) \\ d \text{ is the disparity (horizontal motion)}} [I_L(x'+d, y') - I_R(x', y')]^2$

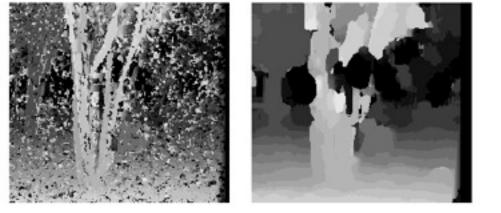


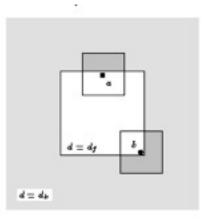
How big should the neighborhood be?

Slide credit: Rick Szeliski

Neighborhood size

Smaller neighborhood: more details Larger neighborhood: fewer isolated mistakes





w = 3

w = 20

Slide credit: Rick Szeliski

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Slide credit: Rick Szeliski

Raw pixel values (correlation)

Slide credit: Rick Szeliski

Raw pixel values (correlation)

Band-pass filtered images [Jones & Malik 92]

Slide credit: Rick Szeliski

Raw pixel values (correlation) Band-pass filtered images [Jones & Malik 92] "Corner" like features [Zhang, ...]

Raw pixel values (correlation) Band-pass filtered images [Jones & Malik 92] "Corner" like features [Zhang, ...] Edges [many people...]

- Raw pixel values (correlation)
- Band-pass filtered images [Jones & Malik 92]
- "Corner" like features [Zhang, ...]
- Edges [many people...]
- Gradients [Seitz 89; Scharstein 94]

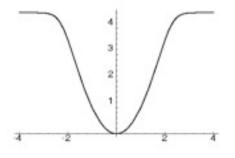
- Raw pixel values (correlation)
- Band-pass filtered images [Jones & Malik 92]
- "Corner" like features [Zhang, ...]
- Edges [many people...]
- Gradients [Seitz 89; Scharstein 94]
- Rank statistics [Zabih & Woodfill 94]

1. For every disparity, compute *raw* matching costs

$$E_0(x, y; d) = \rho(I_L(x' + d, y') - I_R(x', y'))$$

Why use a robust function?

• occlusions, other outliers

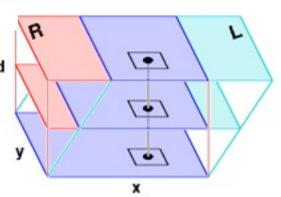


Can also use alternative match criteria

2. Aggregate costs spatially

$$E(x, y; d) = \sum_{(x', y') \in N(x, y)} E_0(x', y', d)$$

 Here, we are using a box filter (efficient moving average implementation)

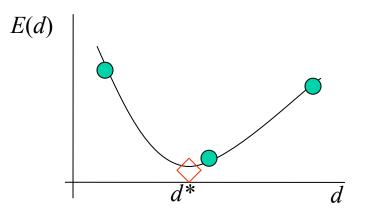


• Can also use weighted average, [non-linear] diffusion...

3. Choose winning disparity at each pixel

$$d(x,y) = \arg\min_{d} E(x,y;d)$$

4. Interpolate to *sub-pixel* accuracy



Slide credit: Rick Szeliski

Advantages:

- gives detailed surface estimates
- fast algorithms based on moving averages
- sub-pixel disparity estimates and confidence

Limitations:

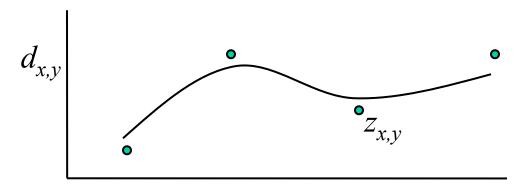
- narrow baseline \Rightarrow noisy estimates
- fails in textureless areas
- gets confused near occlusion boundaries

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Energy minimization

1-D example: approximating splines



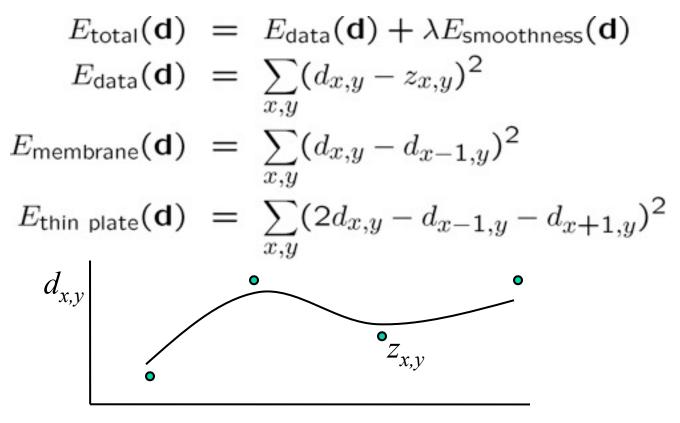
Slide credit: Rick Szeliski

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Energy minimization

1-D example: approximating splines



Slide credit: Rick Szeliski

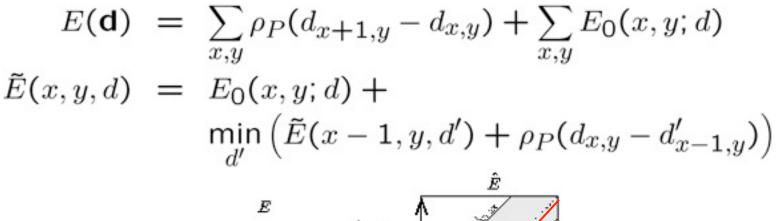
Stereo Topics

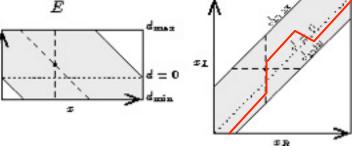
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Evaluate best cumulative cost at each pixel

$$E_{\text{total}}(\mathbf{d}) = E_{\text{data}}(\mathbf{d}) + \lambda E_{\text{smoothness}}(\mathbf{d})$$
$$E_{\text{data}}(\mathbf{d}) = \sum_{x,y} (d_{x,y} - z_{x,y})^2$$
$$E_{\text{smoothness}}(\mathbf{d}) = \sum_{x,y} |d_{x,y} - d_{x-1,y}|$$

1-D cost function





Slide credit: Rick Szeliski

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Disparity space image and min. cost path

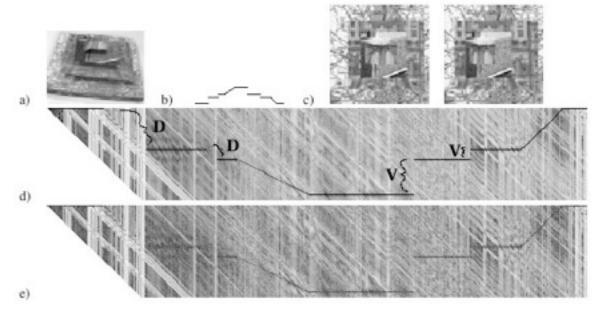


Fig. 4. This figure shows (a) a model of the stereo sloping wedding cake that we will use as a test example, (b) a depth profile through the center of the sloping wedding cake, (c) a simulated, noise-free image pair of the cake, (d) the enhanced, cropped, correlation DSI representation for the image pair in (c), and (e) the enhanced, cropped, correlation DSI for a noisy sloping wedding cake (SNR = 18 dB). In (d), the regions labeled "D" mark diagonal gaps in the matching path caused by regions occluded in the left image. The regions labeled "V" mark vertical jumps in the path caused by regions occluded in the right image.

Slide credit: Rick Szeliski

Sample result (note horizontal streaks)

[Intille & Bobick]





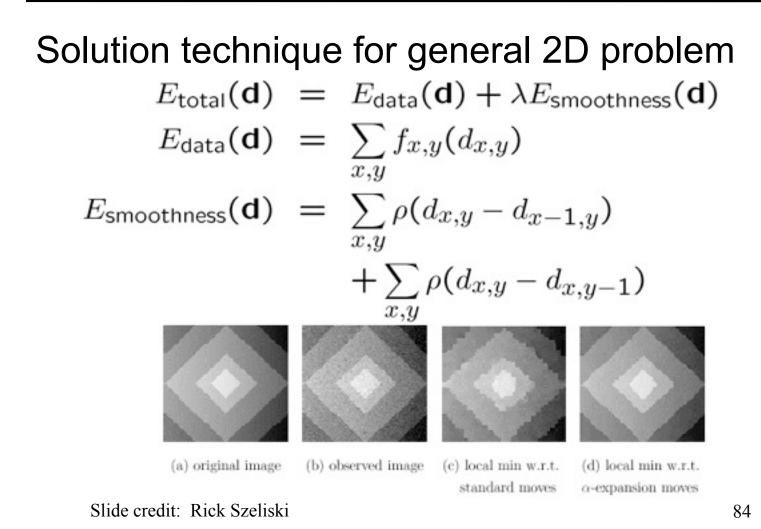
Slide credit: Rick Szeliski

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Stereo Topics

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Graph cuts



graph cuts home page: <u>http://www.cs.cornell.edu/~rdz/graphcuts.html</u>

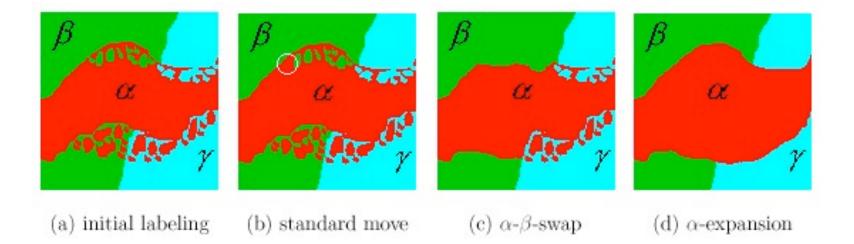
Graph cuts

- α - β swap
- α expansion
- modify smoothness penalty based on edges compute best possible match within integer disparity

graph cuts home page: <u>http://www.cs.cornell.edu/~rdz/graphcuts.html</u>

Graph cuts

Two different kinds of moves:



Slide credit: Rick Szeliski

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Bayesian inference

Formulate as statistical inference problem **Prior model** $p_{P}(\boldsymbol{d})$ Measurement model $p_{M}(\mathbf{I}_{I}, \mathbf{I}_{R} \mid \boldsymbol{d})$ Posterior model $p_M(\boldsymbol{d} \mid \boldsymbol{I}_1, \boldsymbol{I}_R) \propto p_P(\boldsymbol{d}) p_M(\boldsymbol{I}_1, \boldsymbol{I}_R \mid \boldsymbol{d})$ Maximum a Posteriori (MAP estimate): maximize $p_M(\boldsymbol{d} \mid \mathbf{I}_L, \mathbf{I}_R)$

Markov Random Field

Probability distribution on disparity field d(x,y)

$$p_P(d_{x,y}|\mathbf{d}) = p_P(d_{x,y}|\{d_{x',y'}, (x',y') \in \mathcal{N}(x,y)\})$$

$$p_P(\mathbf{d}) = \frac{1}{Z_P} e^{-E_P(\mathbf{d})}$$

$$E_P(\mathbf{d}) = \sum_{x,y} \rho_P(d_{x+1,y} - d_{x,y}) + \rho_P(d_{x,y+1} - d_{x,y})$$

Enforces smoothness or coherence on field

Slide credit: Rick Szeliski

Measurement model

Likelihood of intensity correspondence

$$p_M(I_L, I_R | \mathbf{d}) = \frac{1}{Z_M} e^{-E_0(x,y;d)}$$

$$E_0(x, y; d) = \rho(I_L(x' + d, y') - I_R(x', y'))$$

Corresponds to Gaussian noise for quadratic $\boldsymbol{\rho}$

Slide credit: Rick Szeliski

MAP estimate

Maximize posterior likelihood

$$E(\mathbf{d}) = -\log p(\mathbf{d}|I_L, I_R)$$

$$= \sum_{x,y} \rho_P(d_{x+1,y} - d_{x,y}) + \rho_P(d_{x,y+1} - d_{x,y}) + \sum_{x,y} \rho_M(I_L(x + d_{x,y}, y) - I_R(x, y))$$

Equivalent to *regularization* (energy minimization with smoothness constraints)

Slide credit: Rick Szeliski

Principled way of determining cost function

Principled way of determining cost function Explicit model of noise and prior knowledge

Principled way of determining cost function Explicit model of noise and prior knowledge Admits a wide variety of optimization algorithms:

Principled way of determining cost function Explicit model of noise and prior knowledge Admits a wide variety of optimization algorithms:

• gradient descent (local minimization)

- gradient descent (local minimization)
- stochastic optimization (Gibbs Sampler)

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- mean-field optimization

- gradient descent (local minimization)
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- graph theoretic (actually deterministic) [Zabih]

- gradient descent (local minimization)
- stochastic optimization (Gibbs Sampler)
- mean-field optimization
- graph theoretic (actually deterministic) [Zabih]
- [loopy] belief propagation

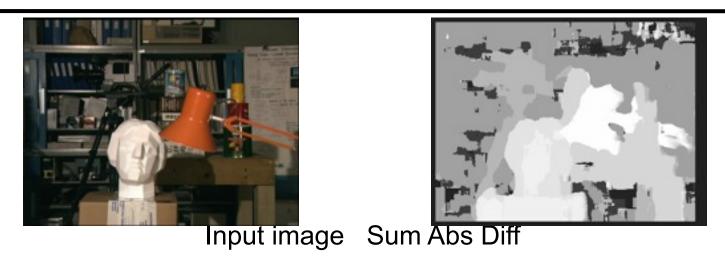
- gradient descent (local minimization)
- stochastic optimization (Gibbs Sampler)
- mean-field optimization
- graph theoretic (actually deterministic) [Zabih]
- [loopy] belief propagation
- large stochastic flips [Swendsen-Wang]



Input image Sum Abs Diff

Mean field Graph cuts

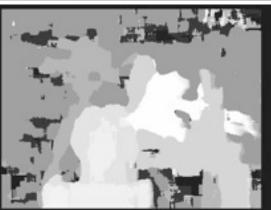
Slide credit: Rick Szeliski



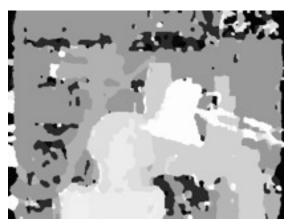
Mean field Graph cuts

Slide credit: Rick Szeliski



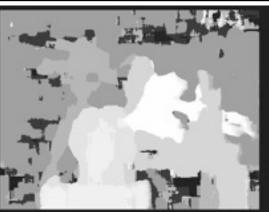


Input image Sum Abs Diff

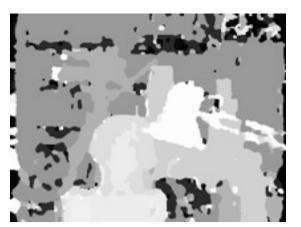


Mean field Graph cuts





Input image Sum Abs Diff





Mean field Graph cuts

Stereo evaluation

vision.middlebury.edu stereo · mview · MRF · flow Stereo Evaluation • Datasets • Code • Submit Daniel Scharstein • Richard Szeliski Welcome to the Middlebury Stereo Vision Page, formerly located at www.middlebury.edu/stereo. This website accompanies our taxonomy and comparison of two-frame stereo correspondence algorithms [1]. It contains: An on-line evaluation of current algorithms · Many stereo datasets with ground-truth disparities Our stereo correspondence software An on-line submission script that allows you to evaluate your stereo algorithm in our framework How to cite the materials on this website:

We grant permission to use and publish all images and numerical results on this website. If you report performance results, we request that you cite our paper [1]. Instructions on how to cite our datasets are listed on the datasets page. If you want to cite this website, please use the URL "vision.middlebury.edu/stereo/".

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Stereo—best algorithms

Error Threshold = 1			Sort by	nonocc		Sort by all					Sort by disc		
		V				▼					V		
Algorithm	Avg. Rank	Tsukuba ground truth			Venus ground truth			Teddy ground truth			Cones ground truth		
		nonocc		<u>disc</u>	nonocc	all	<u>disc</u>	nonocc		<u>disc</u>	nonocc	all	disc
AdaptingBP [17]	2.8	<u>1.11</u> 0	1.37 3	5.797	<u>0.10</u> 1	0.21 2	1.44 1	4.22 4	7.06 2	11.8 4	2.48 1	7.92 2	7.32
DoubleBP2 [35]	2.9	0.88 1	1.29 1	4.76 1	0.13 3	0.45 5	1.87 5	3.53 2	8.30 3	9.63 1	2.90 3	8.78 8	7.79
DoubleBP [15]	4.9	0.88 2	1.29 2	4.76 2	0.14 5	0.60 13	2.00 7	<u>3.55</u> 3	8.71 5	9.70 2	2.90 4	9.24 11	7.80
SubPixDoubleBP [30]	5.6	1.24 10	1.76 13	5.98 8	0.12 2	0.46 8	1.74 4	3.45 1	8.38 4	10.0 3	2.93 5	8.737	7.91
AdaptOvrSegBP [33]	9.9	1.69 22	2.04 21	5.64 8	0.14 4	0.20 1	1.47 2	7.04 14	11.17	16.4 11	3.60 11	8.96 10	8.84
SymBP+occ [7]	10.8	0.97 4	1.75 12	5.09 4	0.16 8	0.33 3	2.19 8	<u>6.47</u> 8	10.7 8	17.0 14	4.79 24	10.7 21	10.9
PlaneFitBP [32]	10.8	<u>0.97</u> 5	1.83 14	5.26 5	0.17 7	0.51 8	1.71 3	6.65 9	12.1 13	14.7 7	4.17 20	10.7 20	10.6
AdaptDispCalib [36]	11.8	<u>1.19</u> 8	1.42.4	6.15 9	0.23 9	0.34 4	2.50 11	7.80 19	13.6 21	17.3 17	3.62 12	9.33 12	9.72
Segm+visib [4]	12.2	1.30 15	1.57 5	6.92 18	0.79 21	1.06 18	6.76 22	5.00 5	6.54 1	12.3 5	3.72 13	8.62 8	10.2
C-SemiGlob [19]	12.3	2.61 29	3.29 24	9.89 27	0.25 12	0.57 10	3.24 15	<u>5.14</u> 8	11.8 8	13.0 6	2.77 2	8.35 4	8.20
SO+borders [29]	12.8	1.29 14	1.71 9	6.83 15	0.25 13	0.53 9	2.26 9	7.02 13	12.2 14	16.3 9	3.90 15	9.85 18	10.2
DistinctSM [27]	14.1	<u>1.21</u> 9	1.75 11	6.39 11	0.35 14	0.69 18	2.63 13	7.45 18	13.0 17	18.1 19	3.91 18	9.91 18	8.32
CostAggr+occ [39]	14.3	1.38 17	1.96 17	7.14 19	0.44 18	1.13 19	4.87 19	6.80 11	11.9 10	17.3 10	3.60 10	8.57 6	9.36

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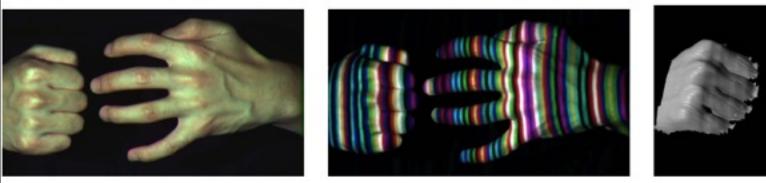
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S

Stereo Topics

- Special, simple system, main idea
- More general camera conditions, epipolar constraints
 - epipolar geometry
 - epipolar algebra
- Image rectification
- Stereo matching (likelihood term)
- Stereo regularization (prior term)
- Inference
 - dynamic programming
 - graph cuts
- Structured light

Active stereo with structured light



Li Zhang's one-shot stereo

Project "structured" light patterns onto the object

• simplifies the correspondence problem

Li Zhang, Brian Curless, and Steven M. Seitz. Rapid Shape Acquisition Using Color Structured Light and Multi-pass Dynamic Programming. In *Proceedings of the 1st International Symposium on 3D Data Processing, Visualization, and Transmission (3DPVT)*, Padova, Italy, June 19-21, 2002, pp. 24-36.

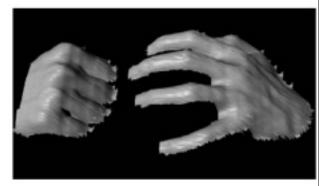
Slide credit: Rick Szeliski

96

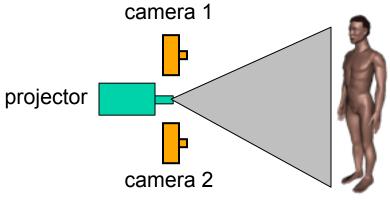
Active stereo with structured light







Li Zhang's one-shot stereo



Project "structured" light patterns onto the object

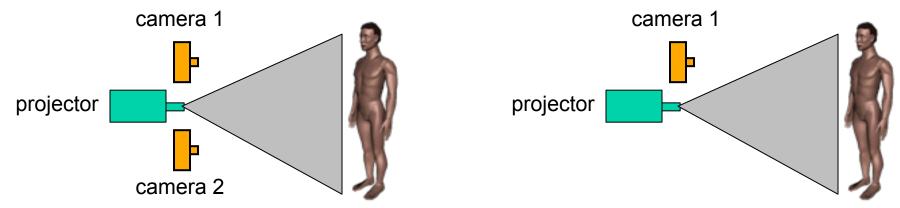
• simplifies the correspondence problem

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Active stereo with structured light



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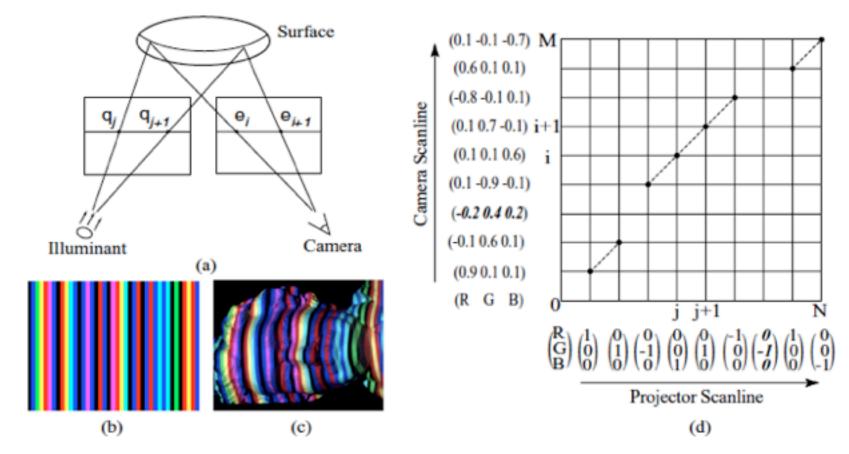


Figure 2. Summary of the one-shot method. (a) In optical triangulation, an illumination pattern is projected onto an object and the eflected light is captured by a camera. The 3D point is reconstructed from the relative displacement of a point in the pattern and mage. If the image planes are rectified as shown, the displacement is purely horizontal (one-dimensional). (b) An example of he projected stripe pattern and (c) an image captured by the camera. (d) The grid used for multi-hypothesis code matching. The horizontal axis represents the projected color transition sequence and the vertical axis represents the detected edge sequence, both taken for one projector and rectified camera scanline pair. A match represents a path from left to right in the grid. Each vertical axis, and q_j , the color transition vectors shown below the horizontal axis. The score for the entire match is the summation of scores along its path. We use dynamic programming to find the optimal path. In the illustration, the camera edge in bold italics corresponds to a false detection, and the projector edge in bold italics is missed due to, e.g., occlusion.

Li Zhang, Brian Curless, and Steven M. Seitz

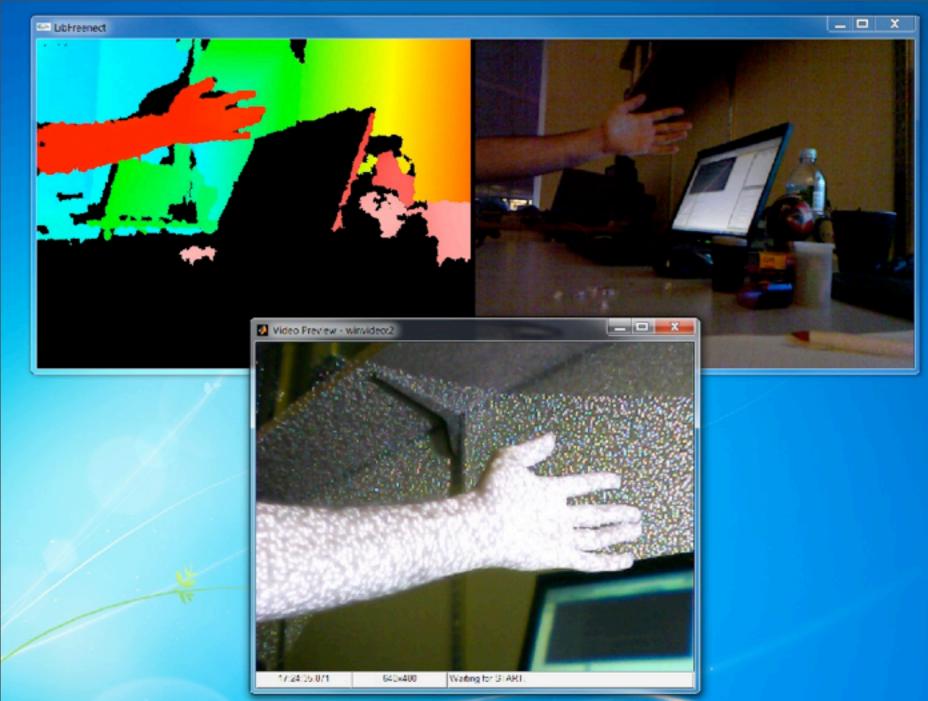








Monday, March 14, 2011

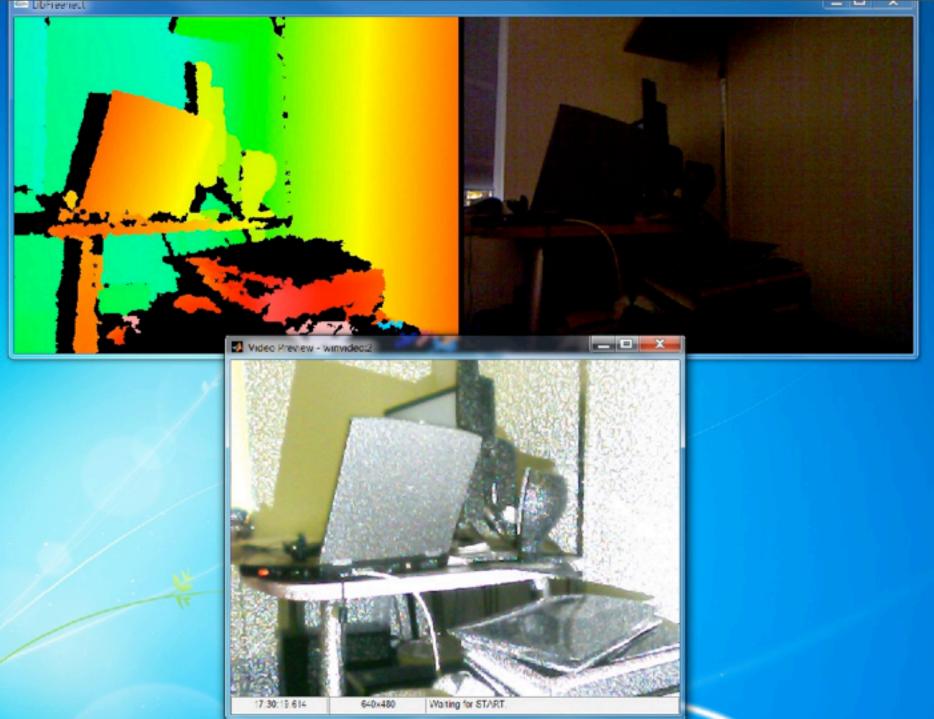




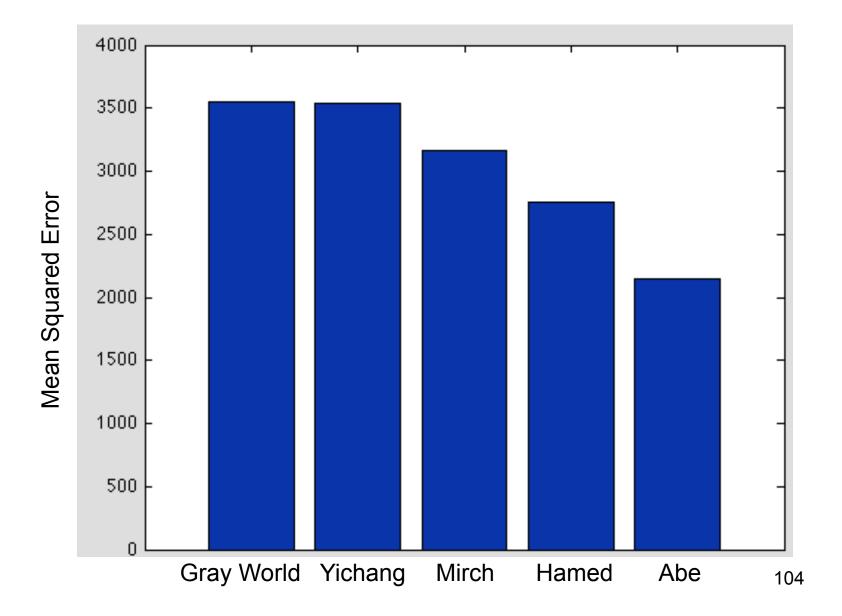








OSA Illumination Spectrum Estimation Contest Leaders, March 13, 2011



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